

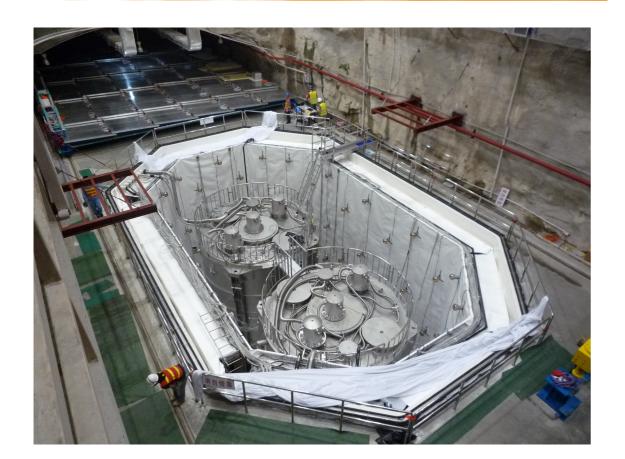
Recent Results From Daya Bay

Chao Zhang



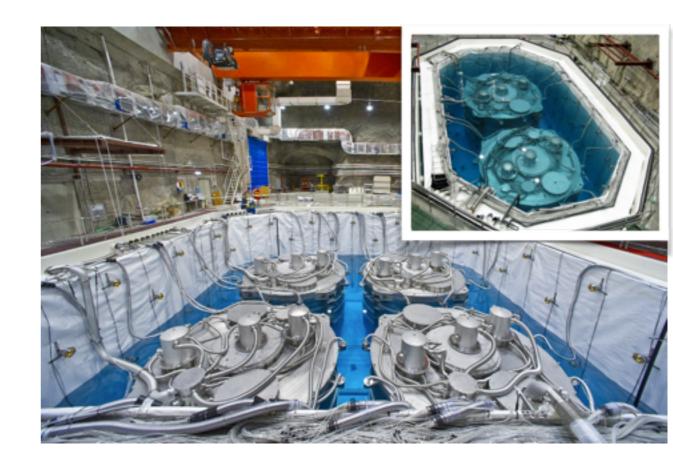
INFO 11 (Jul 18)

- Smooth progress
 - Two ADs in Daya Bay Near Hall installed
 - Dry-run test completed
 - Muon system for Daya Bay Near Hall completed. Water pool fill in August
 - Hall 1 physics data taking soon

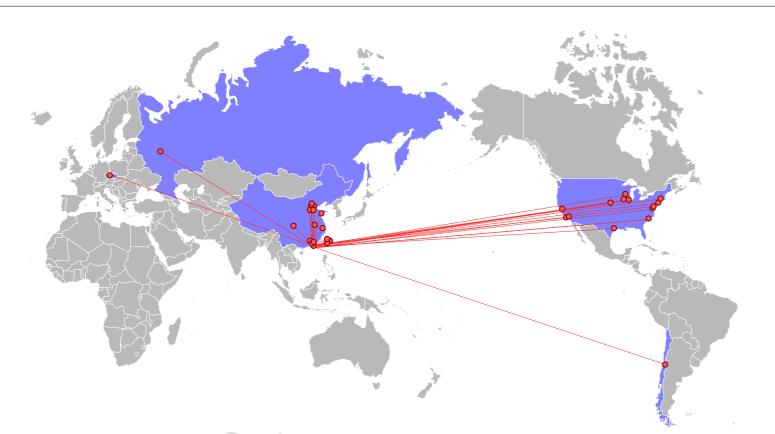


INFO 15 (today)

- Many important results
 - Most precise measurement of sin²2θ₁₃ with 621 days of data
 - Comparable precision of Δm^2_{32} with accelerator experiments
 - Light sterile neutrino search
 - Precise measurements of reactor antineutrino spectrum



The Daya Bay Collaboration





Asia (22)

IHEP, Beijing Normal Univ., Chengdu Univ. of Sci. and Tech., CGNPG, CIAE, Dongguan Univ. of Tech., Nanjing Univ., Nankai Univ., NCEPU, Shandong Univ., Shanghai Jiao tong Univ., Shenzhen Univ., Tsinghua Univ., USTC, Xi'an Jiaotong Univ., Zhongshan Univ., Univ. of Hong Kong, Chinese Univ. of Hong Kong, National Taiwan Univ., National Chiao Tung Univ., National Unive Chongqing Univ.

North America (17)

BNL, LBNL, Iowa State Univ., RPI, Illinois Inst. Tech., Princeton, UC-Berkeley, UCLA, Univ. of Cincinnati, Univ. of Houston, Univ. of Wisconsin, William & Mary, Virginia Tech., Univ. of Illinois-Urbana-Champaign, Siena, Temple Univ, Yale

Europe (2)

JINR, Dubna, Russia; Charles University, Czech Republic

South America (1)

Catholic Univ. of Chile



The Daya Bay Experiment Daya Bay was designed for Far Hall 1615 m from Ling Ao I a sensitivity to 1985 m from Daya Bay $\sin^2 2\theta_{13} < 0.01$ at 90% C.L. 350 m overburden Ling Ao Near Hall 481 m from Ling Ao I 526 m from Ling Ao II 112 m overburden 3 Underground Experimental Halls Entrance - Ling Ao II Cores **Ling Ao I Cores** Daya Bay Near Hall 363 m from Daya Bay ■ 17.4 GW_{th} power 98 m overburden 8 operating detectors ■ 160 t total target mass Daya Bay Cores

Double Chooz, France



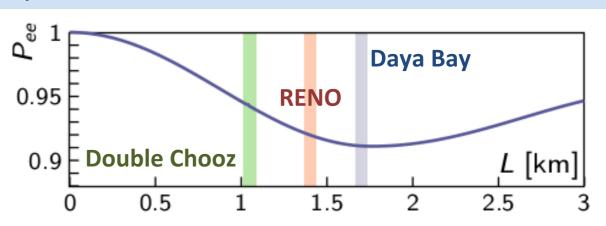
RENO Korea

Daya Bay, China



Baseline Optimization

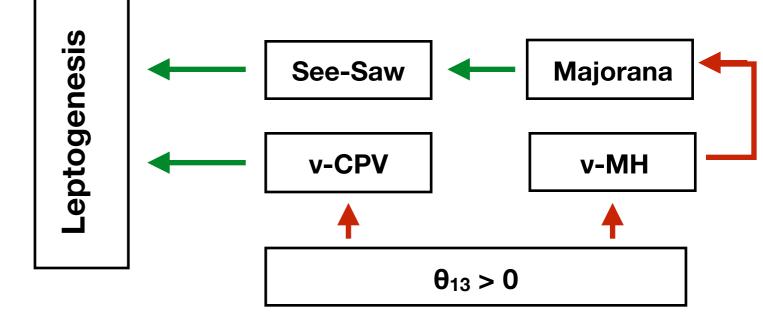
- Detector locations optimized to known parameter space of $|\Delta m^2_{ee}|$
- Far site maximizes term dependent on $\sin^2 2\theta_{13}$



Go strong, big and deep! Depth [m.w.e] Reactor [GW_{th}] Target [tons] **Double Chooz** 8.6 $16(2 \times 8)$ 300, 120 (far, near) $32(2 \times 16)$ **RENO** 16.5 450, 120 $160 (8 \times 20)$ 860, 250 17.4 **Daya Bay** Large Signal Low Background



$$\begin{pmatrix} \nu_e \\ \nu_{\mu} \\ \nu_{\tau} \end{pmatrix} = U_{\text{PMNS}} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$





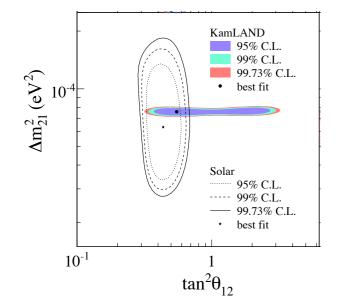
$$\begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \cos \theta_{13} & 0 & \sin \theta_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin \theta_{13} e^{i\delta} & 0 & \cos \theta_{13} \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix}$$

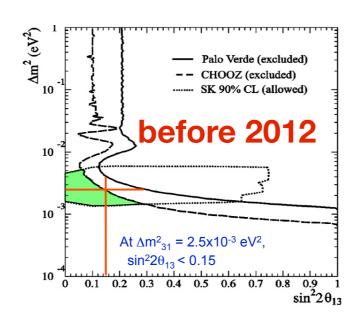
$$\begin{pmatrix}
1 & 0 & 0 \\
0 & \cos \theta_{23} & \sin \theta_{23} \\
0 & -\sin \theta_{23} & \cos \theta_{23}
\end{pmatrix}$$

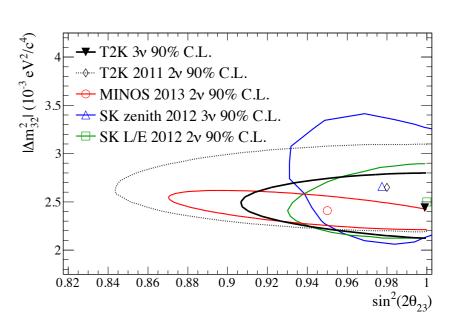
 $\theta_{12} \sim 35^{\circ}$ Solar v Long-Baseline Reactor v

 $\theta_{13} < 10^{\circ}$ Short-Baseline Reactor v Atmospheric v Accelerator v

 $\theta_{23} \sim 45^{\circ}$ Accelerator v

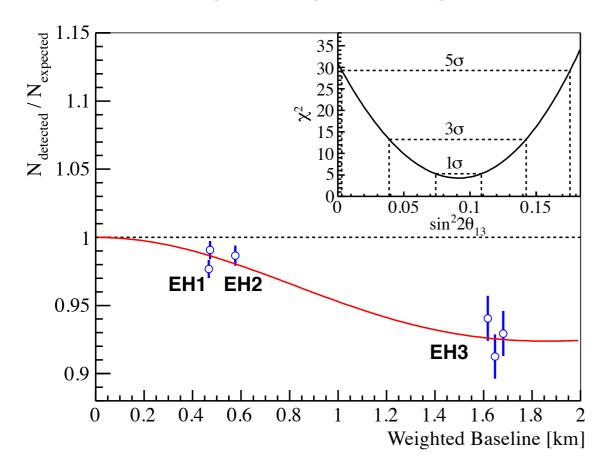






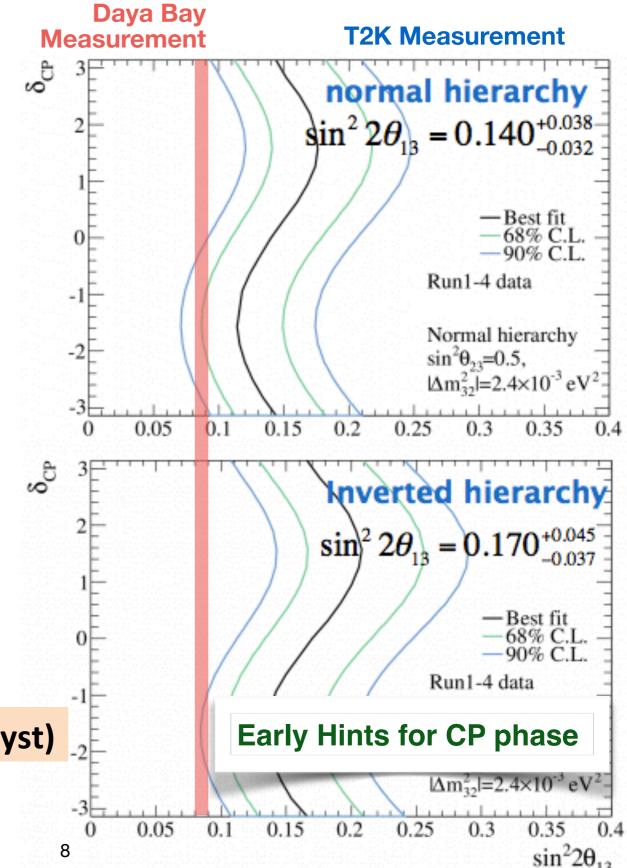
Discovery of non-zero θ13

Daya Bay 55 days





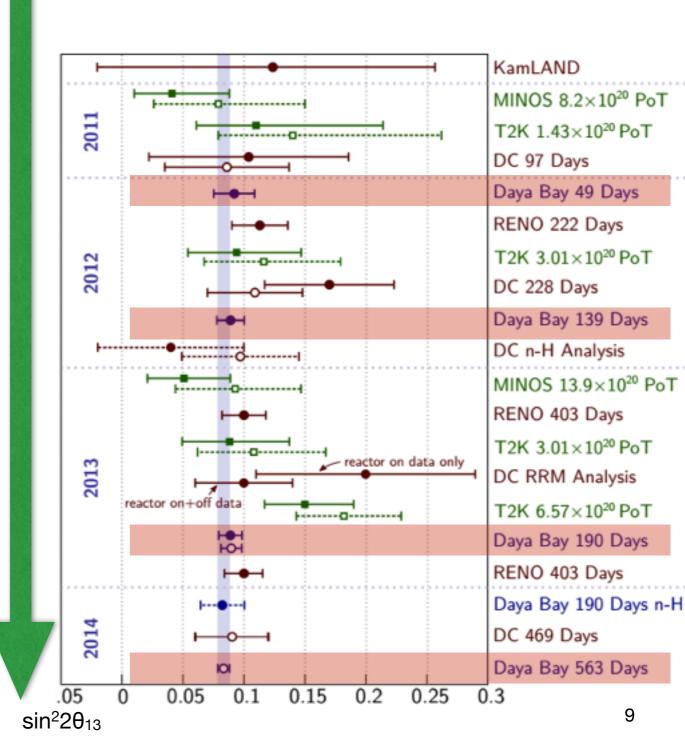
 $\sin^2 2\theta_{13} = 0.092 \pm 0.016 \text{ (stat)} \pm 0.005 \text{ (syst)}$



Scientific Goals After 2012

- Precision measurement of $\sin^2 2\theta_{13}$
 - aim to determine $\sin^2 2\theta_{13}$ to better than 3% by the end of 2017
- Precision measurement of atmospheric mass splitting
 Δm²₃₂ (to better than 3%)
- Search for new neutrino physics such as a light sterile neutrino
- Precision measurement of reactor antineutrino flux and spectrum





Scientific Goals After 2012

- Precision measurement of sin²2θ₁₃
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Relative measurement

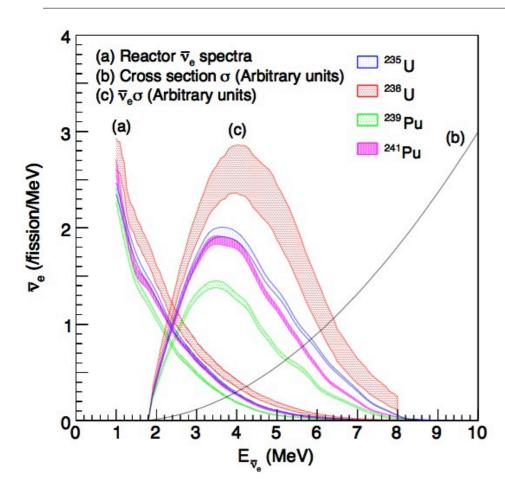
- Reactor-related
 systematics:
 Far/Near relative
 measurement
- Detector-related
 systematics:
 multiple functionally identical detectors (4 Near + 4 Far)

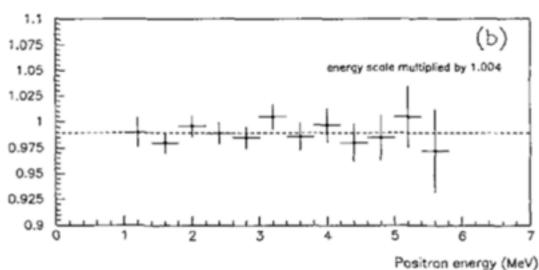
 Precision measurement of reactor antineutrino flux and spectrum

Absolute measurement

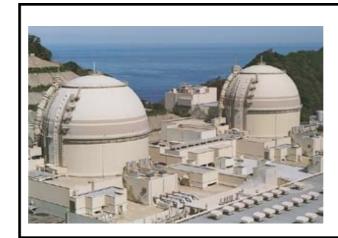
 Extensive calibration and Data / MC comparison

Reactor Antineutrinos





Comparison of Bugey 3 spectrum relative to ILL+Vogel model. *Phys. Lett. B 374, 243 (1996)*



Nuclear Reactor

- pure $\bar{\nu}_e$ source (Free!)
- 6 $\bar{\nu}_e$ / fission
- 6 x 10^{20} $\bar{\nu}_e$ / sec / 3GW_{th}

ILL+Vogel Model

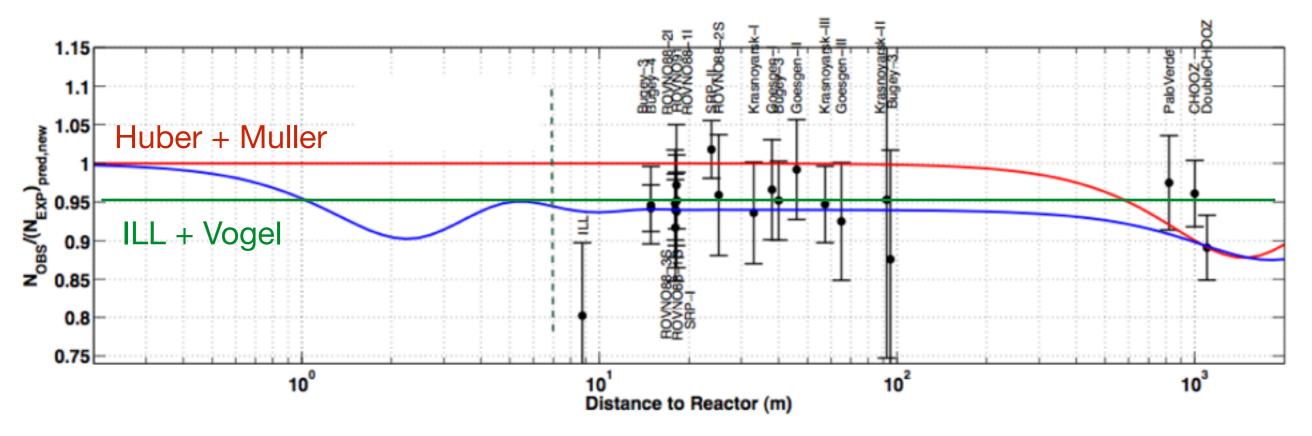
- ²³⁵U, ²³⁹Pu, ²⁴¹Pu neutrino spectra converted from β-spectra measurements at ILL
 - ~ 2.7% uncertainty (conversion from β-spectra to antineutrino spectra)

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Phys. Lett. B160, 325 (1985),
Phys. Lett. B118, 162 (1982)
Phys. Lett. B218, 365 (1989),
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- ²³⁸U neutrino spectrum calculated theoretically
 - ~ 10% uncertainty (but only ~10% of the flux)

Reactor Antineutrino Anomaly

Reanalysis of reactor antineutrino flux in 2011 improved the theoretical uncertainty (to ~2%), but lead to a few % increase in flux prediction that doesn't agree with previous reactor antineutrino flux measurements.



G.Mention et. al, PRD 83, 073006 (2011) P. Huber, PRC 84, 024617 (2011)

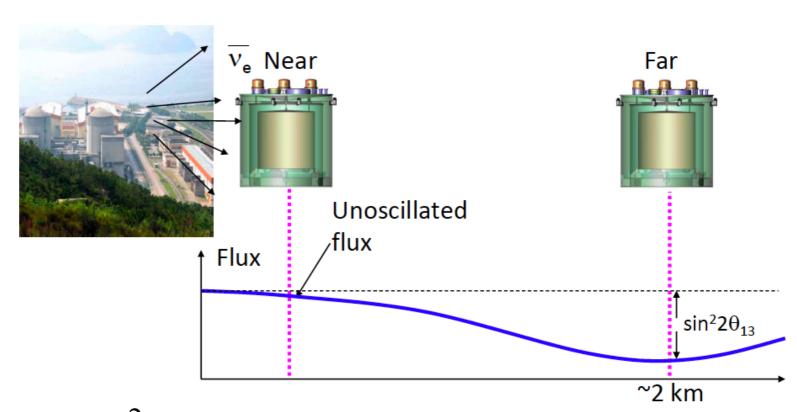
 However, recent re-evaluation of the systematics in the flux calculation suggests that the uncertainty should be > 4% due to forbidden beta decays.

A.C. Hayes et. al, PRL 112, 202501 (2014)

Far/Near Relative Measurement

$$P(\bar{\nu}_e \to \bar{\nu}_e) \simeq 1 - \sin^2 2\theta_{13} \sin^2(\Delta m_{ee}^2 \frac{L}{4E}) - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2(\Delta m_{21}^2 \frac{L}{4E})$$

- Far/Near measurement to largely cancel reactor flux uncertainty (a factor of ~1/20)
- "Functionally Identical detectors" to cancel detector systematics



$$\frac{N_{far}}{N_{near}} = \left(\frac{N_{p,f}}{N_{p,n}}\right) \cdot \left(\frac{L_n}{L_f}\right)^2 \cdot \left(\frac{\varepsilon_f}{\varepsilon_n}\right) \cdot \left(\frac{P_{survival}(E, L_f)}{P_{survival}(E, L_n)}\right)$$

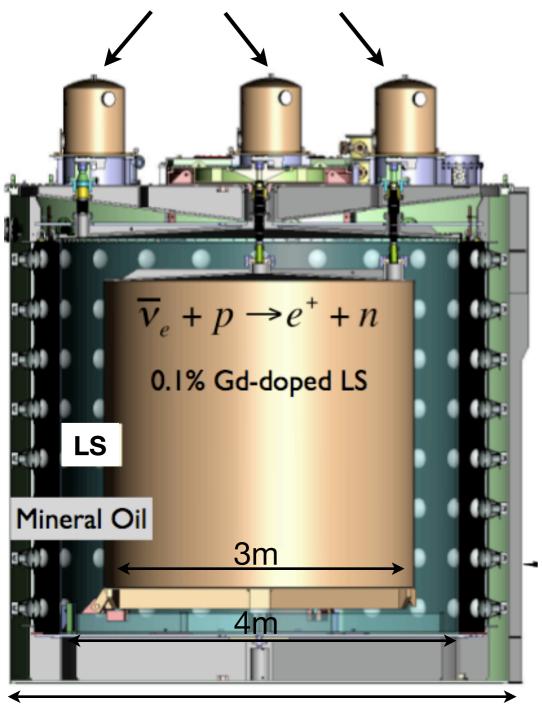
Far/Near Neutrino Ratio Detector Target Mass

Distance from Reactor

Detector Efficiency Survival Probability (θ_{13})

Anti-neutrino Detector (AD)

Automated Calibration Units



5_m

8 functionally identical detectors

$$\frac{N_{\rm f}}{N_{\rm n}} = \left(\left(\frac{N_{\rm p,f}}{N_{\rm p,n}}\right) \left(\frac{L_{\rm n}}{L_{\rm f}}\right)^2 \left(\frac{\epsilon_{\rm f}}{\epsilon_{\rm n}}\right) \left[\frac{P_{\rm sur}(E,L_{\rm f})}{P_{\rm sur}(E,L_{\rm n})}\right]$$

Each detector has 3 nested cylindrical zones separated by Acrylic Vessels:

Inner: 20 tons Gd-doped LS (target volume)

Mid: 20 tons LS (gamma catcher)

Outer: 40 tons mineral oil (buffer)

Each detector has:

192 8-inch Photomultipliers (PMTs)

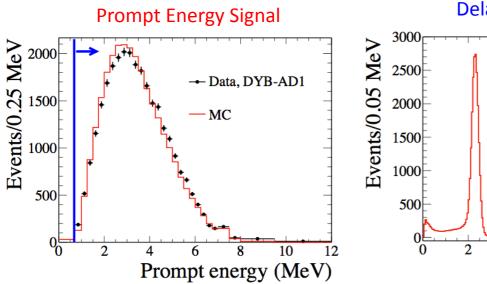
Optical reflectors at top/bottom of cylinder

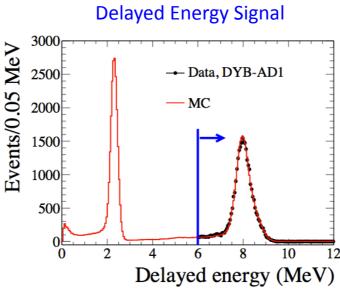
- effectively 12% photocoverage
- ~ 160 photoelectrons / MeV
- ~ 8%/√E (MeV) energy resolution

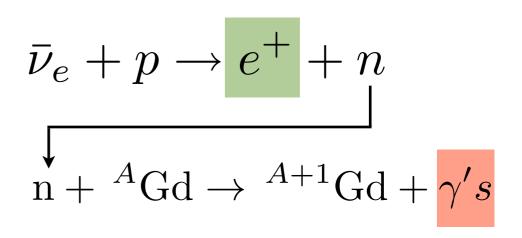
Anti-neutrino Detection Method

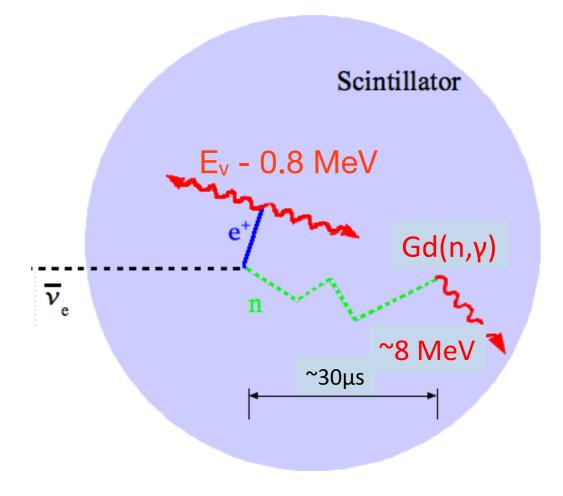
Inverse Beta Decay (IBD)

- E_{threshold} = 1.8 MeV
- 'Large' cross section σ~10⁻⁴² cm²
- Distinctive coincidence signature in a large liquid scintillator detector









Gd-LS defines the target volume. Fiducial volume cut is not necessary.

Muon Veto System

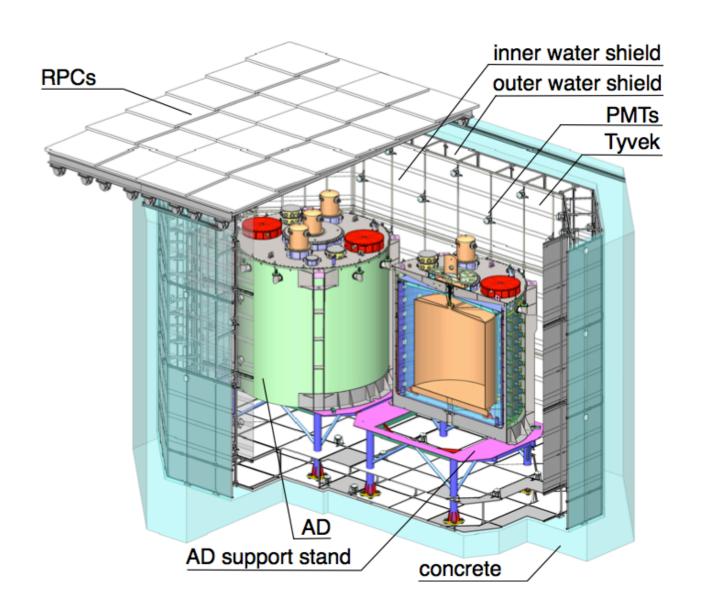
Multiple muon veto detectors >2.5m thick two-sector active water shield and RPC

Water Cherenkov

- Detectors submerged in water shielded against external neutrons and gammas
- Optically separated with Tyvek sheets into inner / outer region for better muon tracking
- 8-inch PMTs mounted on frames, 288 @Near, 384 @Far

Resistive Plate Chamber (RPC)

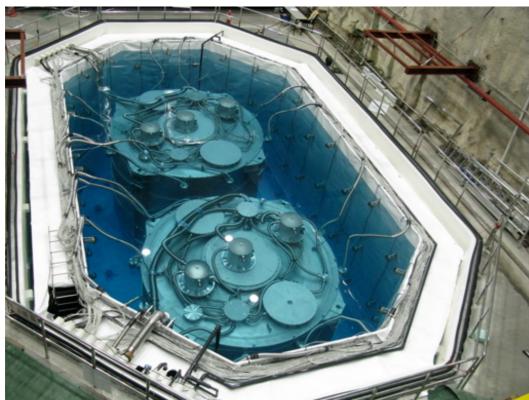
- Independent muon tagging
- Retractable roof above pool
- 54 modules @Near, 81 @Far

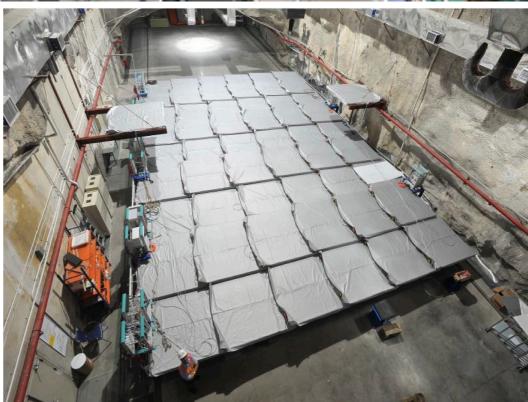


Antineutrino Detector Installation

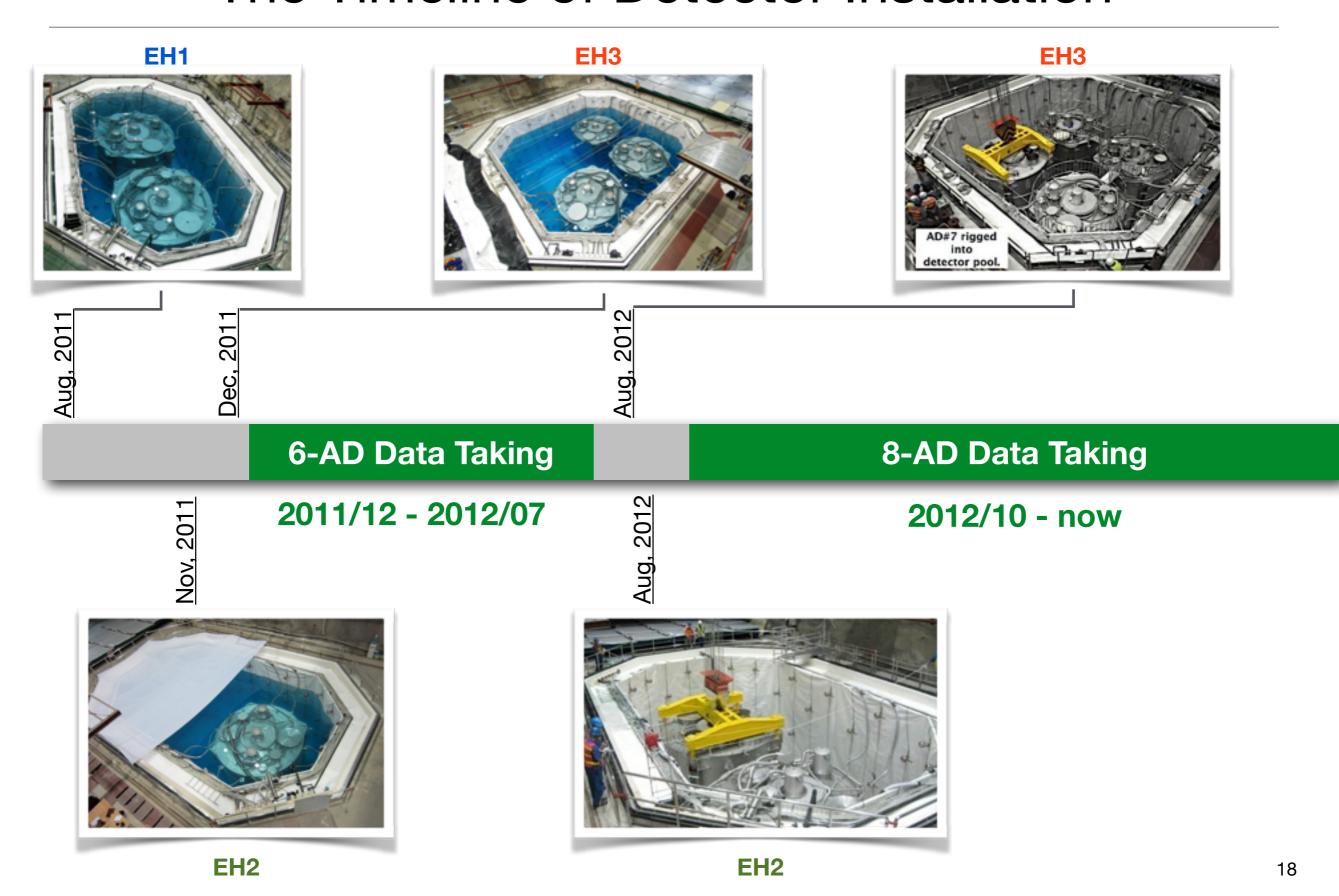








The Timeline of Detector Installation



Analysis Results

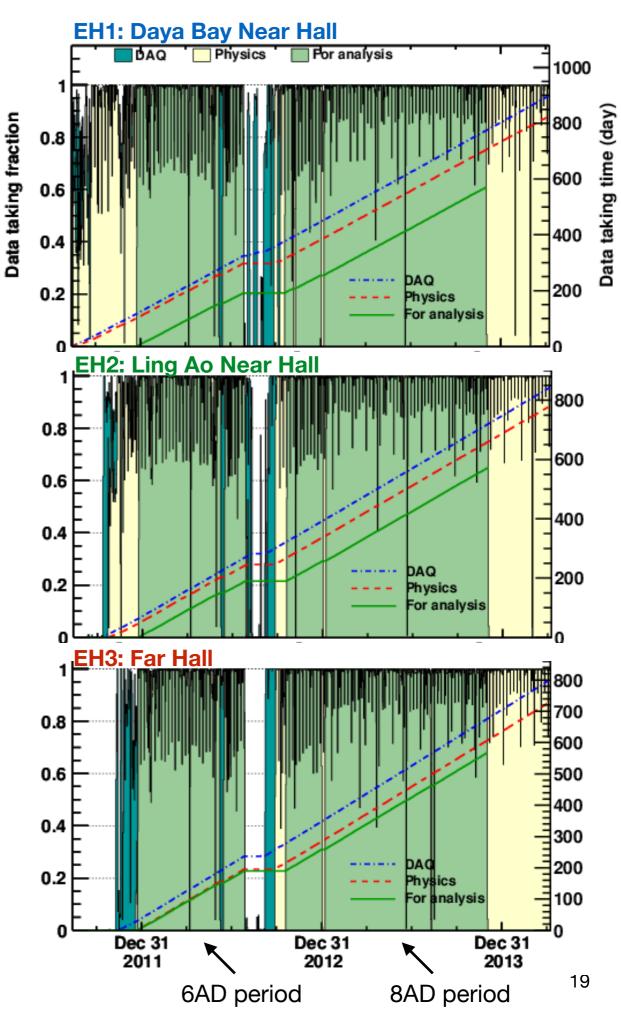
<u>Using Combined 6+8AD period (621 days):</u>

- Oscillation analysis: $\sin^2 2\theta_{13}$ and Δm^2_{ee}
 - 4 times more statistics than our previously published results (*PRL* 112, 061801 (2014))

Using 6-AD period (217 days):

- Independent measurement of sin²2θ₁₃ using neutron capture on hydrogen
- Light sterile neutrino search
- Measurement of reactor antineutrino flux and spectrum

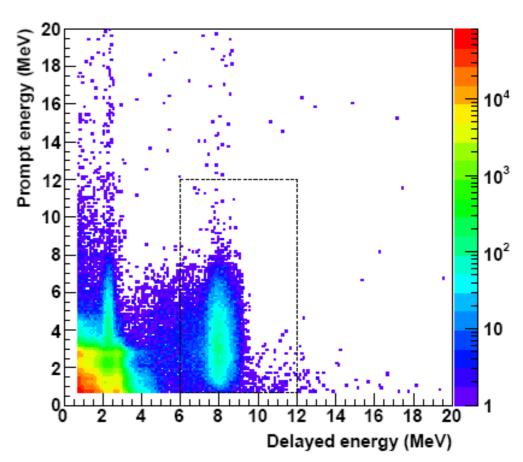
6-AD period: 2011/12/24 - 2012/07/28 (217 days) 8-AD period: 2012/10/19 - 2013/11/27 (404 days)

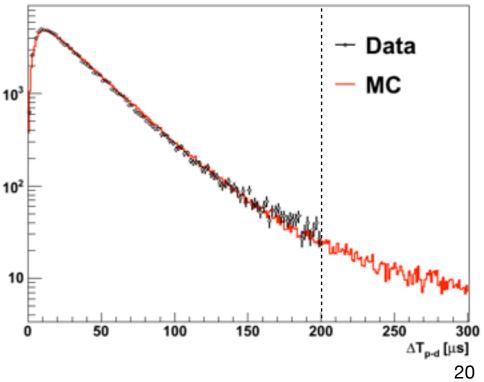


Antineutrino Candidate Selection

- Reject PMT flashers
- Muon veto:
 - Water pool Muon: reject 0.6ms
 - AD Muon (>20 MeV): reject 1 ms
 - AD Shower Muon (>2.5 GeV): reject 1s
- Prompt positron Energy: 0.7 MeV < Ep < 12 MeV
- Delayed neutron Energy: 6 MeV < Ed < 12 MeV
- Neutron Capture time: 1 us < Δt < 200 us
- Multiplicity cut: only select isolated candidate pairs

	Efficiency	Uncertainty	
		Correlated	Uncorrelated
Target Protons		0.47%	0.03%
Flasher cut	99.98%	0.01%	0.01%
Delayed Energy cut	92.7%	0.97%	0.12%
Prompt Energy cut	99.81%	0.10%	0.01%
Capture time cut	98.70%	0.12%	0.01%
Gd capture ratio	84.2%	0.95%	0.10%
Spill-in correction	104.9%	1.50%	0.02%
Combined	80.6%	2.1%	0.2%

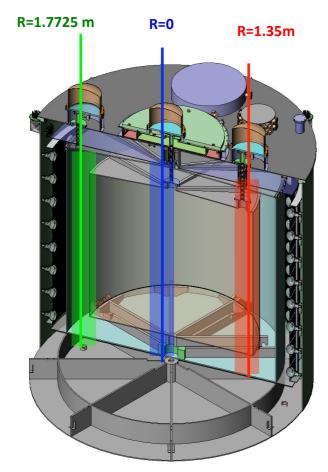




Calibration System

3 Automated Calibration Units (ACUs) per detector

 3 sources for each z axis on a turntable (position accuracy < 7 mm)



- 10 Hz ⁶⁸Ge (2 x 0.511 MeV γ's)
- 100 Hz ⁶⁰Co gamma source (1.173 + 1.332 MeV γ's) + 0.7 Hz ²⁴¹Am¹³C neutron source (3.5 MeV n without γ)
- LED diffuser ball for PMT gain and timing

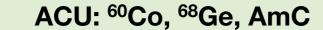


Manual Calibration System (one-time)

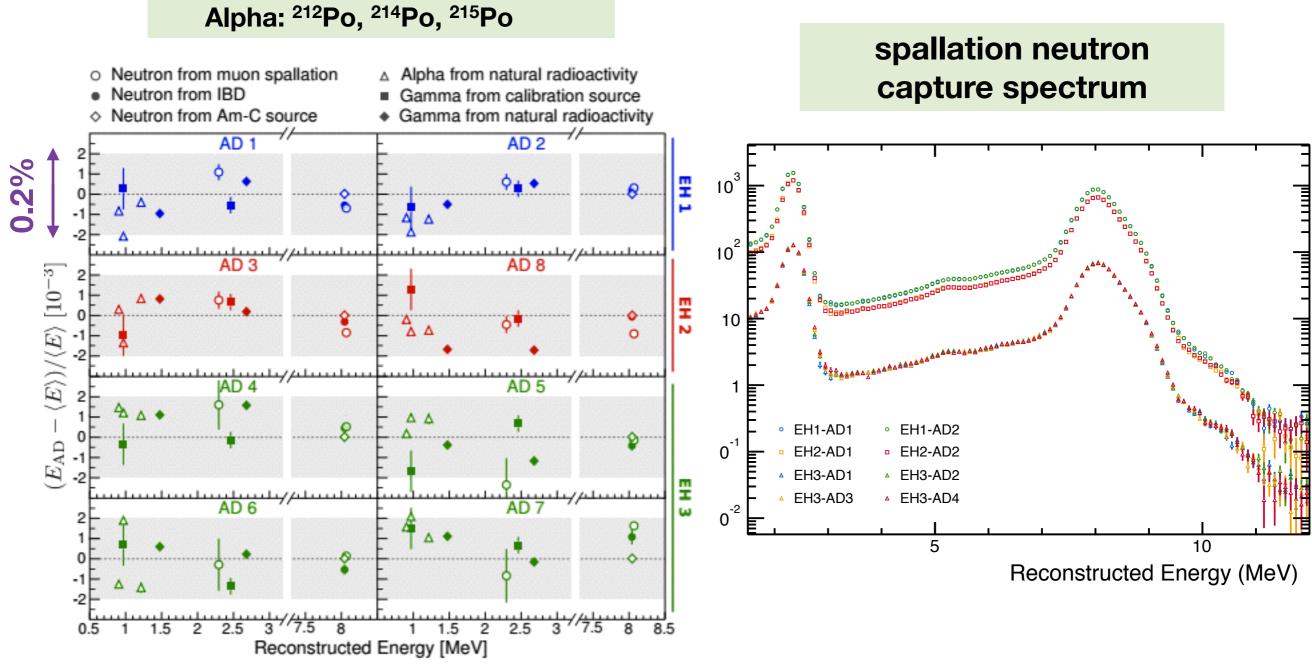
- MCS installed on AD#1 in summer 2012
- 60Co + ²³⁹Pu¹³C composite source
- 4π deployment

- Simultaneous, fully-automated weekly deployment for all 8 ADs
- Special calibration campaign during summer 2012 with temporary sources
 - ¹³⁷Cs, ⁵⁴Mn, ⁴⁰K, ²⁴¹Am⁹Be, ²³⁹Pu¹³C
- Also have methods to calibrate in-situ
 - PMT gains: dark noise
 - Energy (light-yield): spallation neutron

Relative Energy Scale

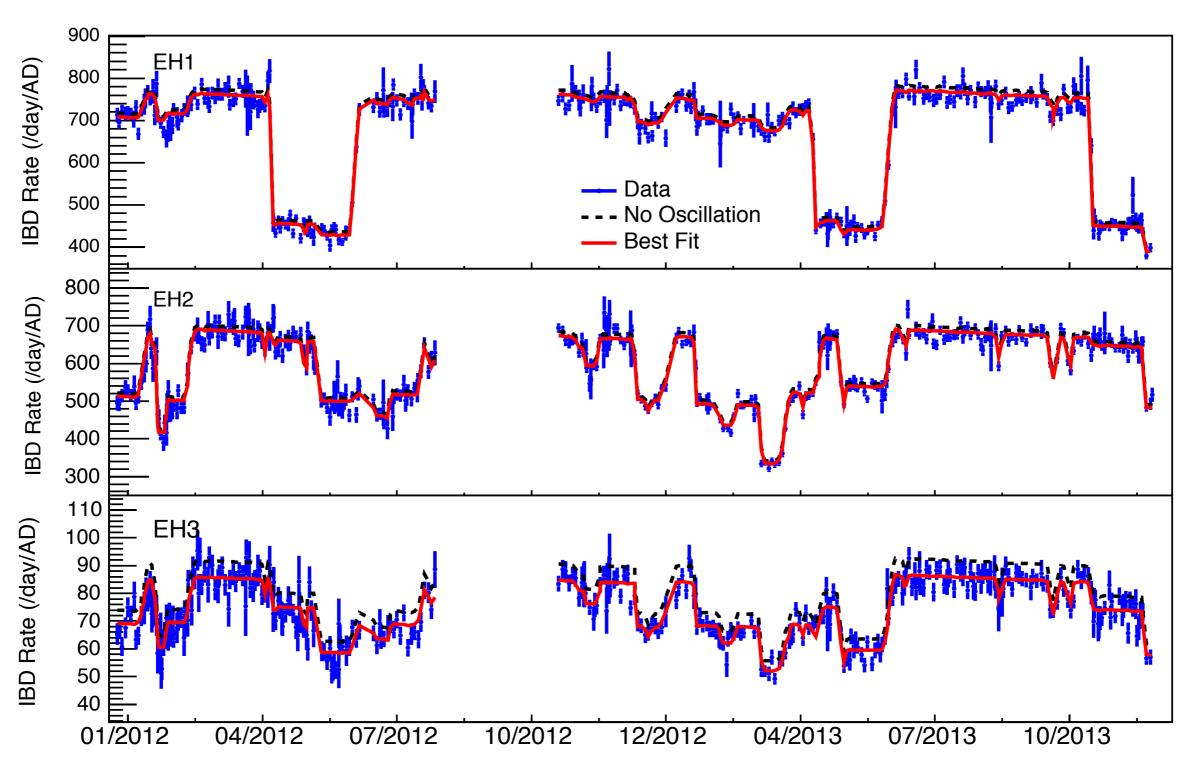


Spallation: nGd, nH Gamma: 40K, 208TI



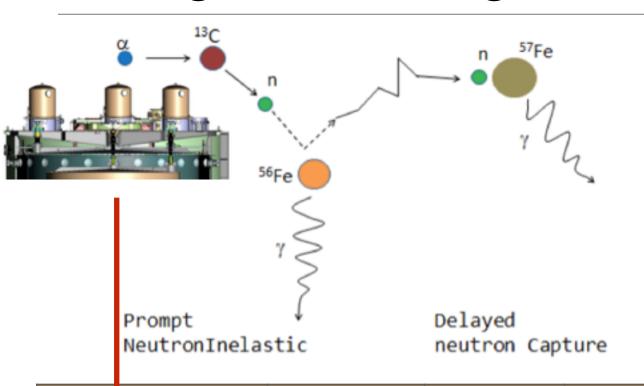
< 0.2% variation in reconstructed energy between ADs

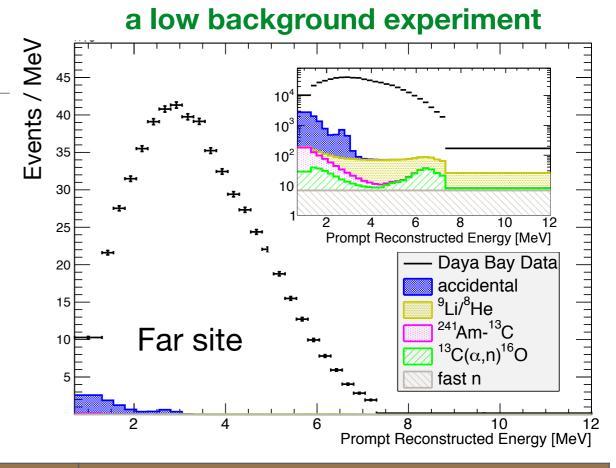
Over 1 million antineutrino interactions! (150k at the far site)



Detected rate strongly correlated with reactor flux

Background Budget

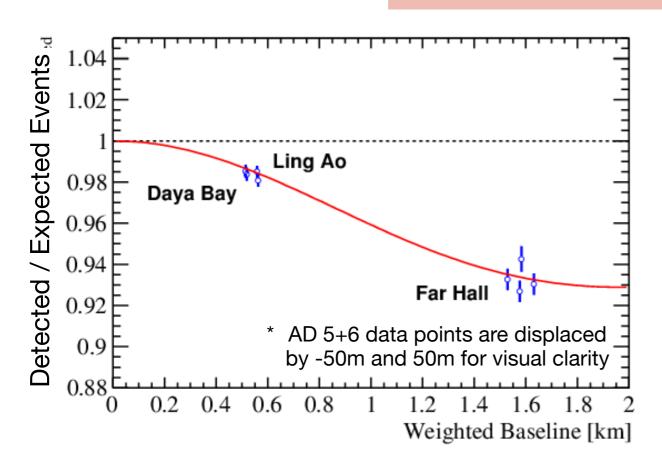




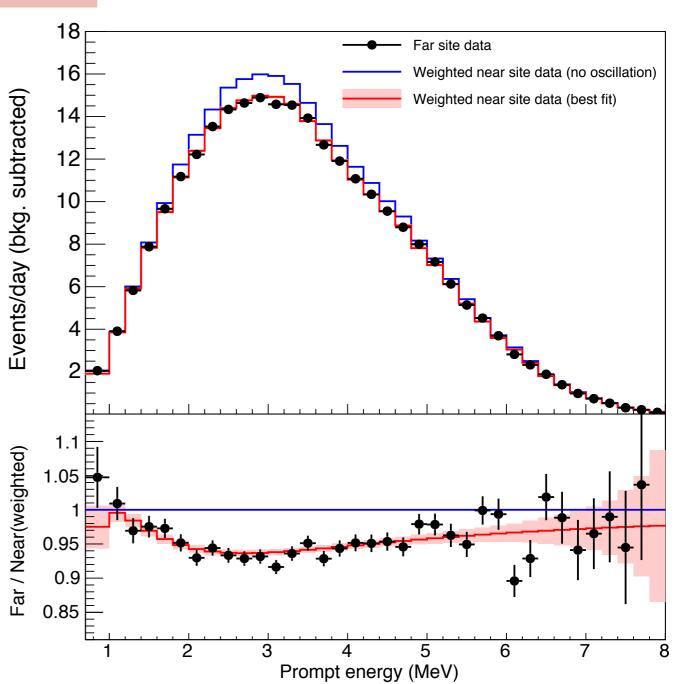
Background	Near	Far	Uncertainty	Method
Accidentals	1.4%	2.3%	negligible	statistically calculated from uncorrelated singles
AmC source	0.03%	0.2%	~50%	MC benchmarked with single gamma and strong AmC source
Li-9 / He-8	0.4%	0.4%	~50%	measured with after-muon events
Fast neutron	0.1%	0.1%	~30%	measured from AD/water/RPC tagged muon events
Alpha-n	0.01%	0.1%	~50%	calculated from measured radioactivity

Far v.s. Near Comparison

$$P(\bar{\nu}_e \to \bar{\nu}_e) \simeq 1 - \sin^2 2\theta_{13} \sin^2(\Delta m_{ee}^2 \frac{L}{4E}) - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2(\Delta m_{21}^2 \frac{L}{4E})$$



The observed relative rate deficit and relative spectrum distortion are highly consistent with oscillation interpretation



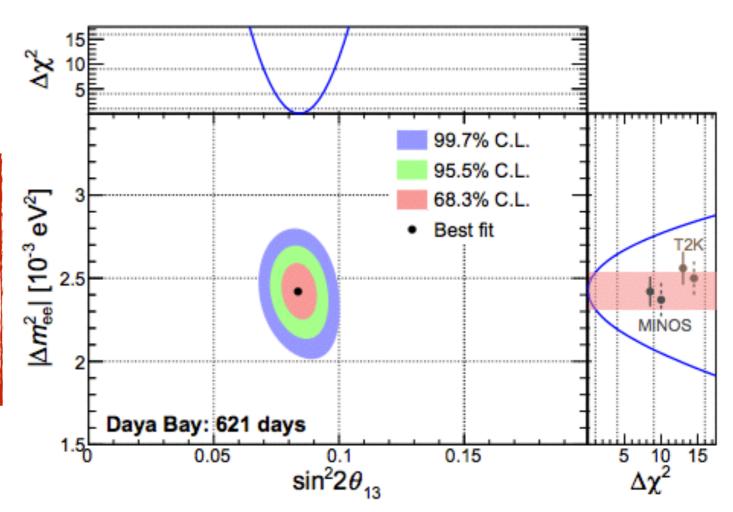
Oscillation Results

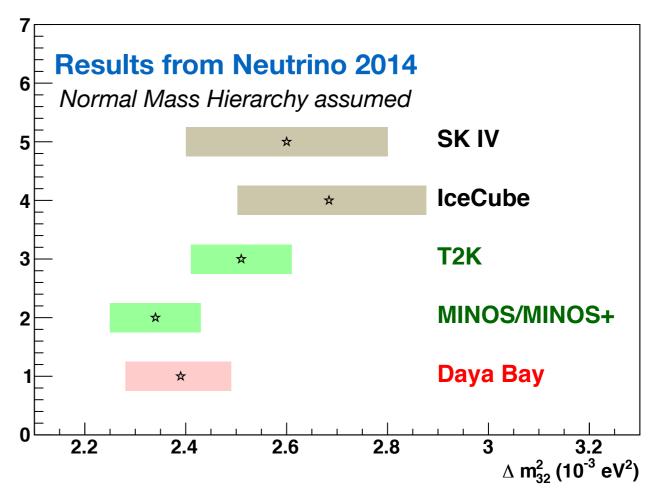
$$\sin^2 2\theta_{13} = 0.084 \pm 0.005$$

 $|\Delta m_{ee}^2| = (2.42 \pm 0.11) \times 10^{-3} \,\text{eV}^2$
 $\chi^2/\text{NDF} = 134.6/146$

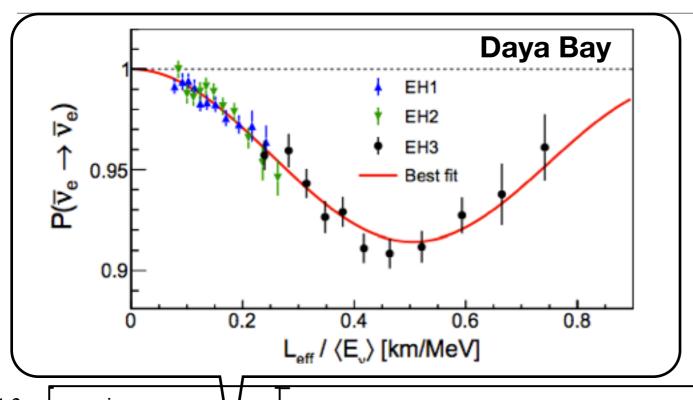
- Most precise measurement of sin²2θ₁₃, precision reached < 6%
- Most precise measurement of atmospheric mass splitting in the electron neutrino disappearance channel
 - consistent and of comparable preciseness with the muon neutrino disappearance experiments

"A new measurement of antineutrino oscillation with the full detector configuration at Daya Bay" arXiv:1505.03456, accepted by PRL

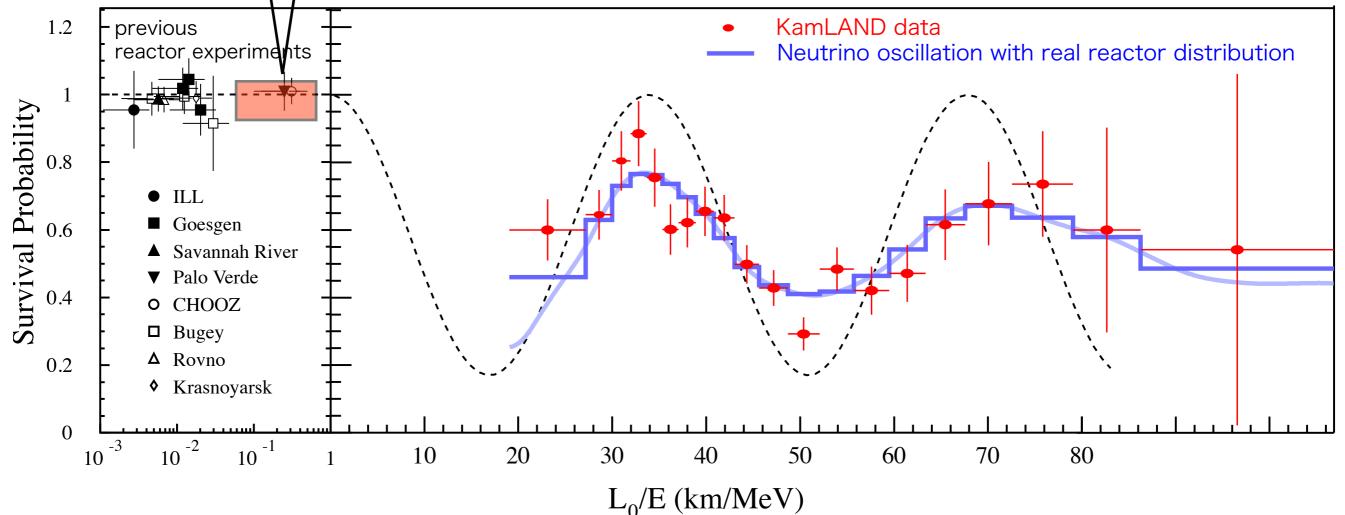




L/E Oscillation in Daya Bay



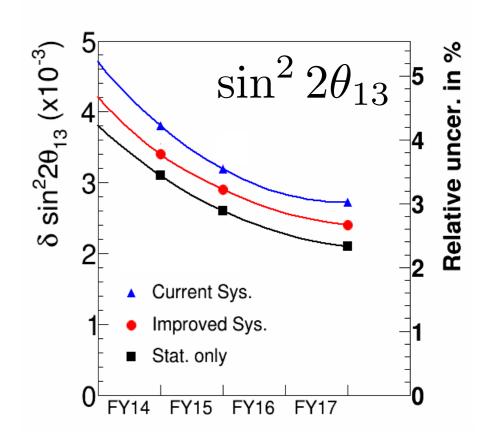
Both atmospheric scale and solar scale oscillations are beautifully demonstrated by reactor neutrino experiments!

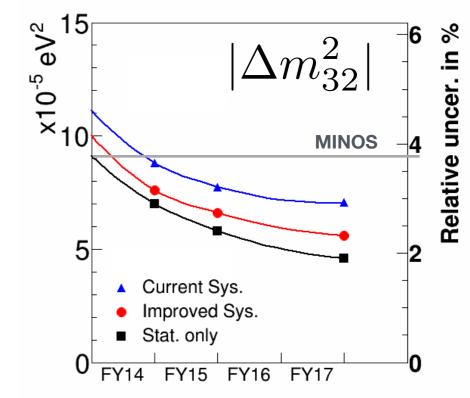


Projected Sensitivity in Oscillation Parameters

Operations through 2017

- Further improve world's most precise measurement of sin²2θ₁₃
 - Enhance the CPV sensitivity for DUNE
 - Test unitarity of PMNS matrix
- Further improve the precision measurement of atmosphere mass splitting in the electron neutrino channel
 - competitive with and complementary to accelerator based experiments, testing 3-flavor model
 - potential hints for neutrino mass hierarchy

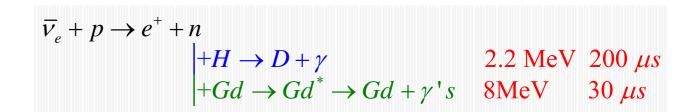


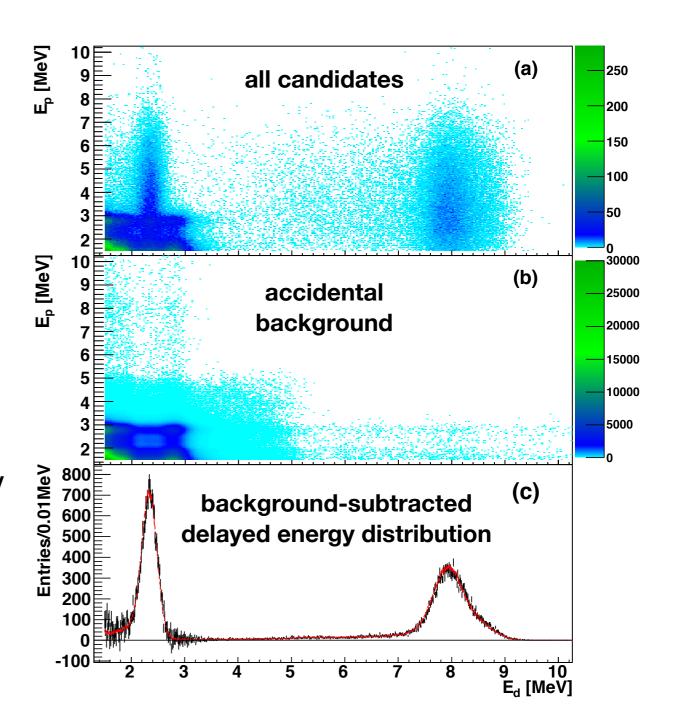


Independent sin²20₁₃ measurement through nH

Advantage

- High statistics (15% capture in the 20ton Gd-LS region and 100% in the 20ton LS region)
- Different systematic uncertainties from nGd analysis
- Challenge
 - High accidental background
 - longer capture time
 - lower delayed energy
- Strategy
 - Raise prompt energy cut Ep > 1.5 MeV
 - Require prompt to delayed distance
 ΔR < 0.5 m
 - Relative measurement to reduce systematics



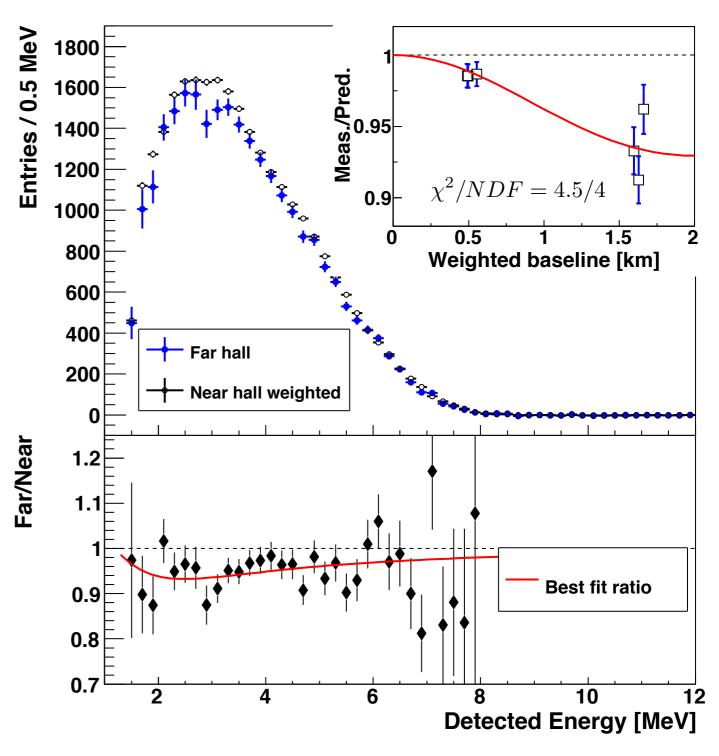


nH Analysis Results

- All 217 days of 6-AD period
- Observed significant rate deficit at far site, <u>rate analysis</u> measures:

$$\sin^2 2\theta_{13} = 0.083 + -0.018$$

- an independent and consistent result with nGd analysis
- another precise measurement of $sin^2 2\theta_{13}$
- Spectrum distortion is consistent with oscillation explanation
 - spectral analysis in progress



Independent θ_{13} measurement with nH, PRD 90, 071101(R) (2014)

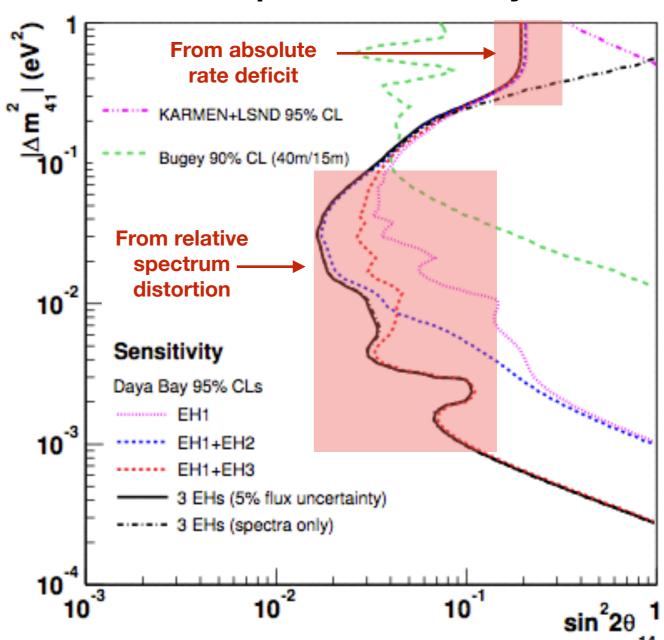
Light Sterile Neutrino Search

- Daya Bay has a unique combination of multiple baselines: EH1 (~350m), EH2 (~500m), EH3 (~1600m)
 - Sterile neutrinos will cause additional spectrum difference between different sites

$$P(ar{
u}_e
ightarrow ar{
u}_e) \simeq 1 - \cos^4 heta_{14}\sin^22 heta_{13}\sin^2\left(rac{\Delta m_{ee}^2L}{4E_
u}
ight)
onumber \ -\sin^22 heta_{14}\sin^2\left(rac{\Delta m_{41}^2L}{4E_
u}
ight)$$

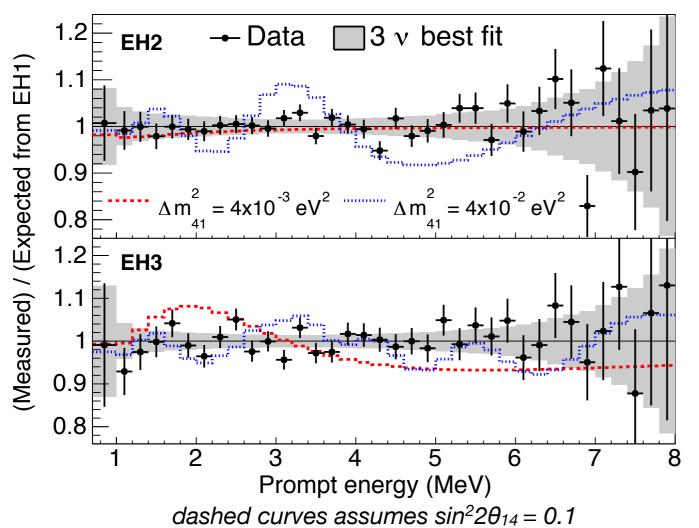
- High sensitivity in the largely unexplored region $\Delta m^2_{41} < 0.1 \text{ eV}^2$
- A robust relative measurement independent of reactor related uncertainties

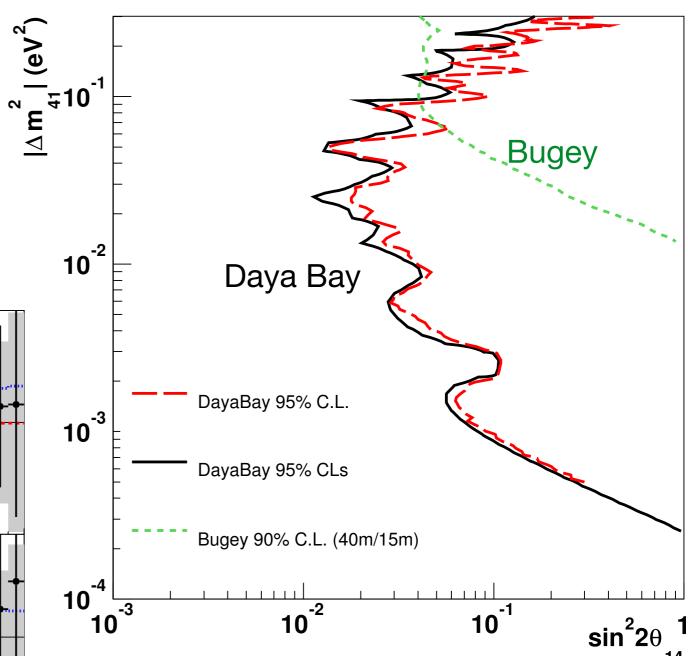
Expected Sensitivity



Light Sterile Neutrino Search Results

- All 217 days of 6-AD period
- Consistent with standard 3-flavor neutrino oscillation model
- Able to set stringent limits in the previously unexplored region $10^{-3} \text{ eV}^2 < \Delta m^2_{41} < 0.1 \text{ eV}^2$





Search for light sterile neutrinos, PRL 113, 141802 (2014)

Reactor Antineutrino Flux and Spectrum

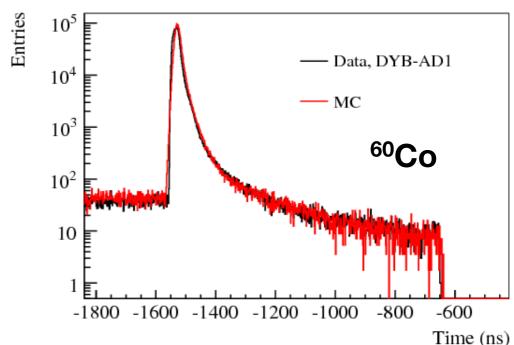
- Integrated Flux: absolute detection efficiency
 - Data /MC comparison

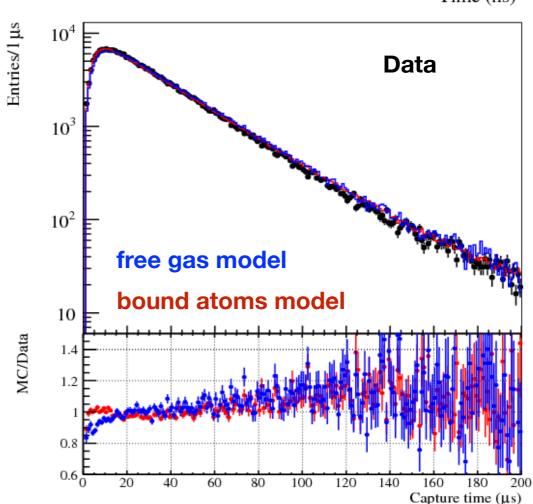
	Efficiency	Uncertainty	
		Correlated	Uncorrelated
Target Protons		0.47%	0.03%
Flasher cut	99.98%	0.01%	0.01%
Delayed Energy cut	92.7%	0.97%	0.12%
Prompt Energy cut	99.81%	0.10%	0.01%
Capture time cut	98.70%	0.12%	0.01%
Gd capture ratio	84.2%	0.95%	0.10%
Spill-in correction	104.9%	1.50%	0.02%
Combined	80.6%	2.1%	0.2%

- Antineutrino Energy Spectrum: absolute energy scale
 - Extensive calibration
 - Energy model building

Daya Bay Monte Carlo

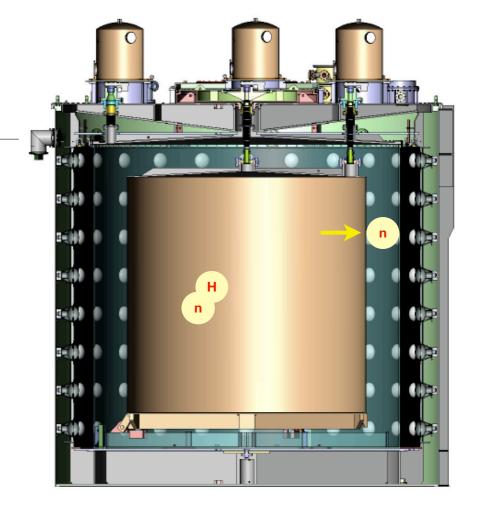
- Full-detector Geant4 simulation implemented in an offline Gaudi-based framework known as NuWa
- Tuning and validation with data
 - Time component of LS
 - Fast (1 ns) 70%, Medium (26 ns) 25.5%, Slow (200 ns) 4.5%
 - Improved thermal neutron scattering simulation
 - free gas model v.s. bound atoms



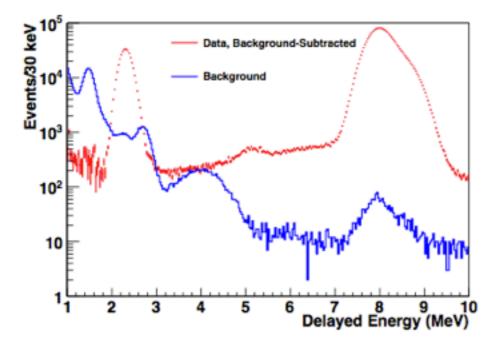


Gd-Capture Ratio

	Efficiency	Uncertainty	
		Correlated	Uncorrelated
Target Protons		0.47%	0.03%
Flasher cut	99.98%	0.01%	0.01%
Delayed Energy cut	92.7%	0.97%	0.12%
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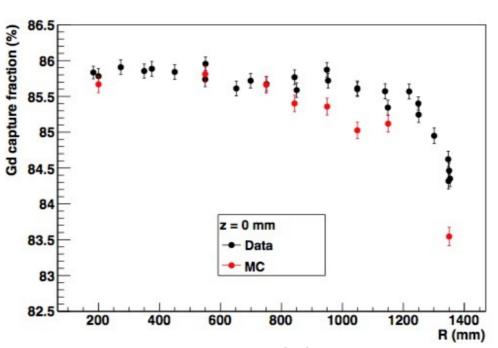


 IBD neutron can capture on other target such as Hydrogen



measured with center neutron source

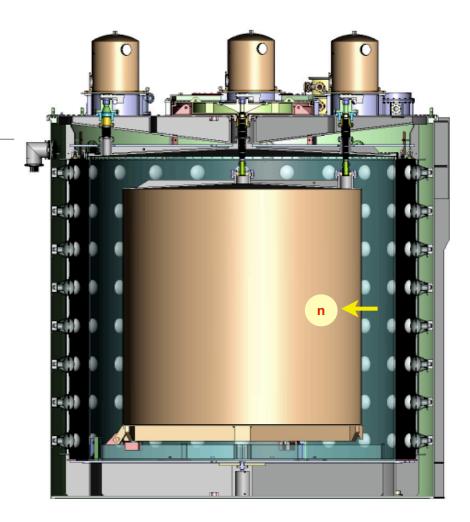
 IBD neutron can leave the GdLS region and capture somewhere else (spill-out)



measured with MCS neutron source

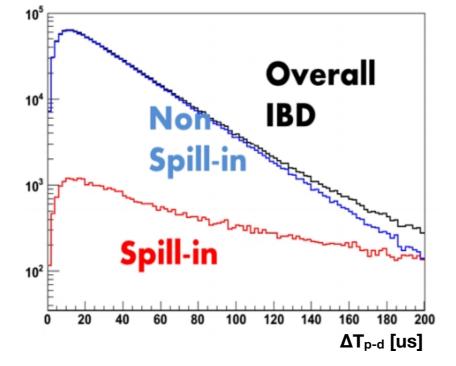
Spill-in

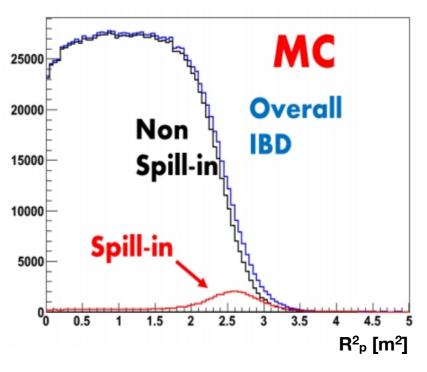
	Efficiency	Uncertainty	
		Correlated	Uncorrelated
Target Protons		0.47%	0.03%
Flasher cut	99.98%	0.01%	0.01%
Delayed Energy cut	92.7%	0.97%	0.12%
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Capture time cut	98.70%	0.12%	0.01%
Gd capture ratio	84.2%	0.95%	0.10%
Spill-in correction	104.9%	1.50%	0.02%
Combined	80.6%	2.1%	0.2%



IBDs can happen outside of GdLS, but the neutron leaks in and capture on Gd

(spill-in)

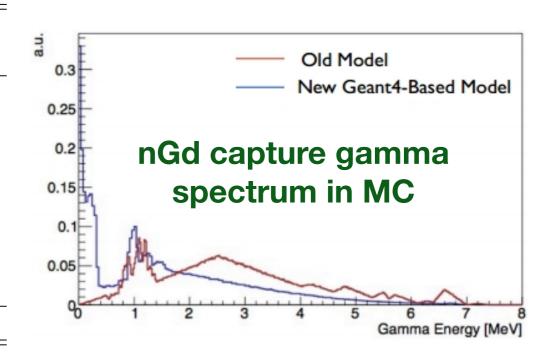




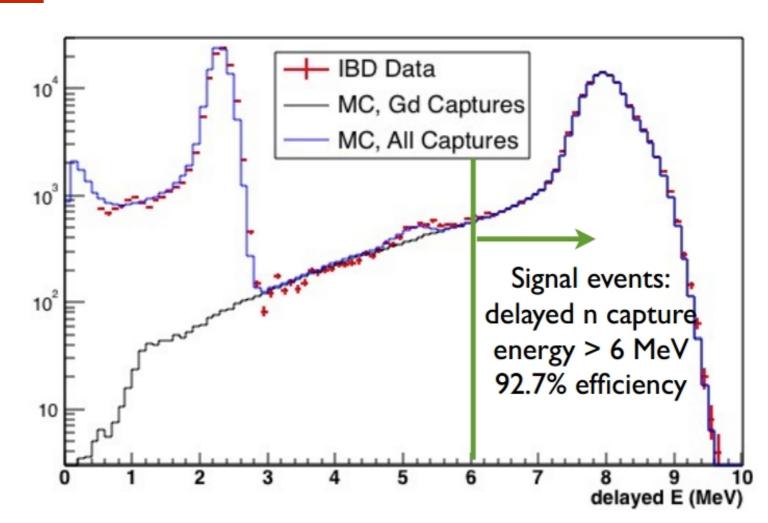
Predicted by MC, benchmarked by IBD time / vertex distribution, and off-center neutron source in LS

Delayed Energy Cut

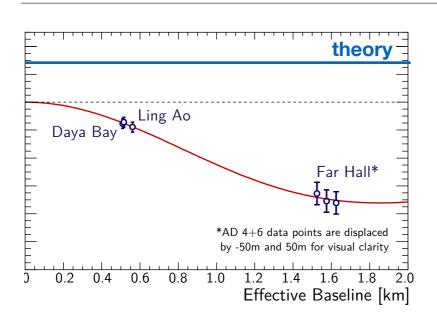
	Efficiency	Uncertainty			
		Correlated	Uncorrelated		
Target Protons		0.47%	0.03%		
Flasher cut	99.98%	0.01%	0.01%		
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Spill-in correction	104.9%	1.50%	0.02%		
Combined	80.6%	2.1%	0.2%		



- The Cascading gammas from nGd Capture can leak out of the scintillator region and cause a long low energy tail
 - measure the tail
 distribution and validate
 with different input nGd
 gamma spectrum models



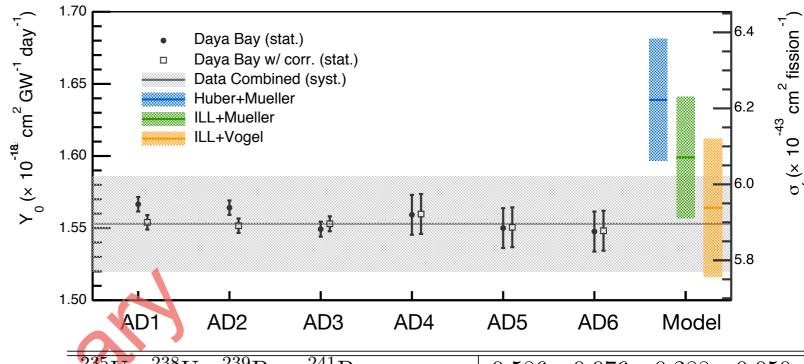
Integrated Reactor Antineutrino Flux

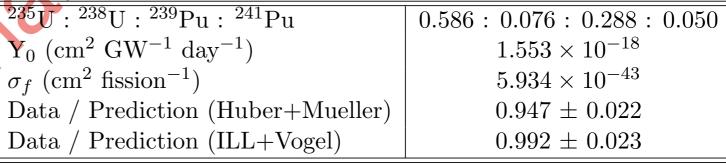


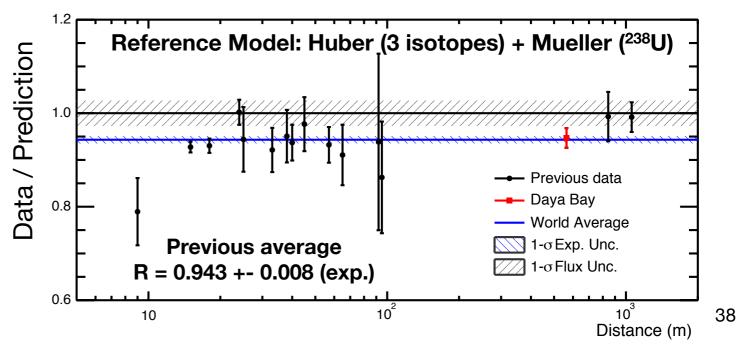
Flux Measurement Uncertainty

	Uncertainty
statistics	0.2%
$ heta_{13}$	0.2%
reactor	0.9%
detector efficiency	2.1%
Total	2.3%

Daya Bay's reactor flux measurement is consistent with previous short baseline experiments





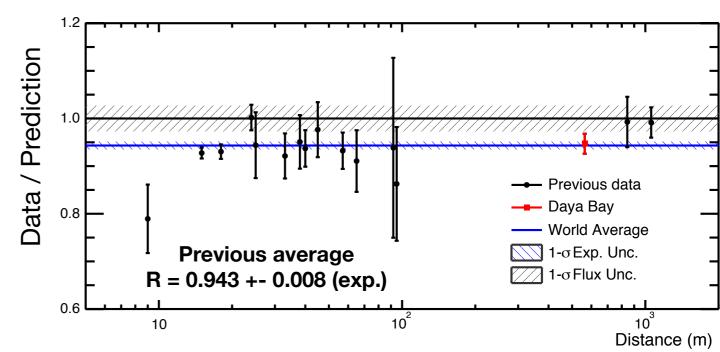


About Flux Measurement Result

- Interpretation one: Daya Bay saw 5.3% deficit with respect to the Huber+Muller model, so there must be a sterile neutrino.
 - Wrong! The reactor flux models could have large uncertainties. The theoretical flux could be over-estimated.
- Interpretation two: The Daya Bay result is consistent with all previous short-baseline measurements, so there is no "anomaly" from experimental point of view.
 - Correct! Our "modern" experiment has validated the results from experiments in 80-90s.
- Question: So is the "reactor antineutrino anomaly" real or not?
 - We don't know yet. To definitively answer that question, one has to observe the spectral "oscillation" feature in the data (like the Daya Bay L/E plot). We need the future ~5m baseline reactor/source experiments to tell us.

One Note About Global Average

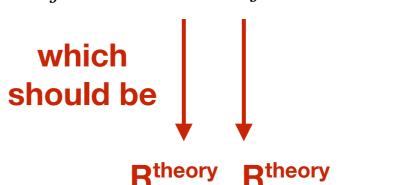
- We obtained R = 0.943+- 0.008 (exp.)
- Many literatures report this number to be 0.928 (1.5% lower).
 - A tricky statistical mistake. They used the measured values to build the theoretical covariance matrix.



$$\chi^2(R_g^{\text{past}}) = (R_g^{\text{past}} - R_i) \cdot V_{ij}^{-1} (R_g^{\text{past}} - R_j),$$

$$V = V^{\exp} + V^{\text{theory}}$$

$$V_{ij}^{\text{theory}} = R_i^{\text{obs}} R_j^{\text{obs}} (\sigma^{\text{theory}})^2$$



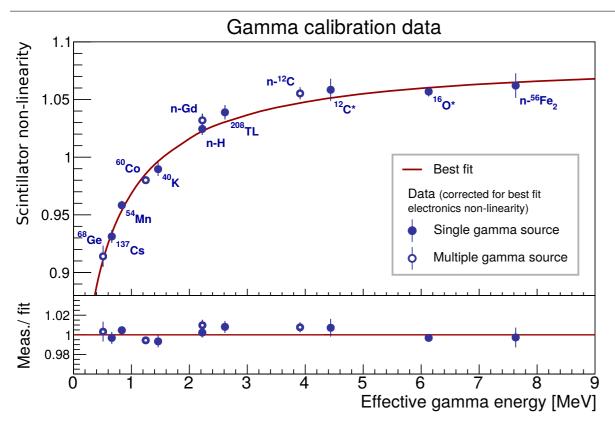
Otherwise could lead to an interesting puzzle that the average is smaller than any of measurements

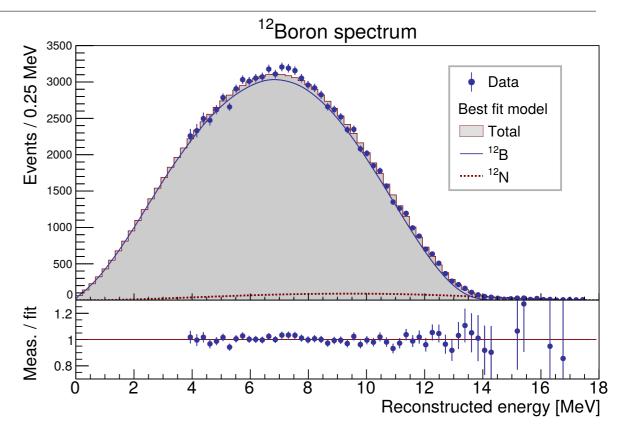
a well known statistics problem originally described in G. D'Agostini, NIMA 346 (1994) 306

see also

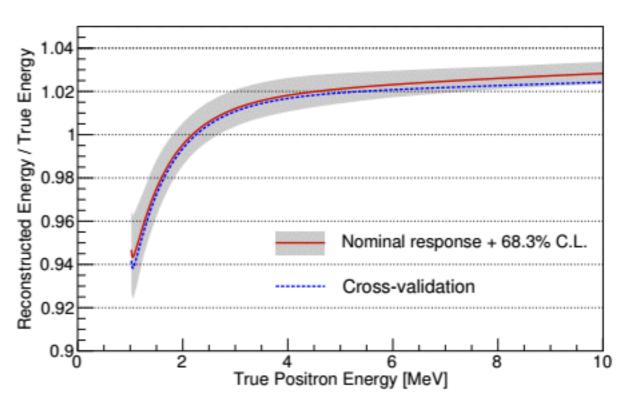
V. Blobel, "Some Comments on \$\chi^2\$ Minimisation Applications,"SLAC-R-0703, pp.101–105.

Energy Nonlinearity Calibration



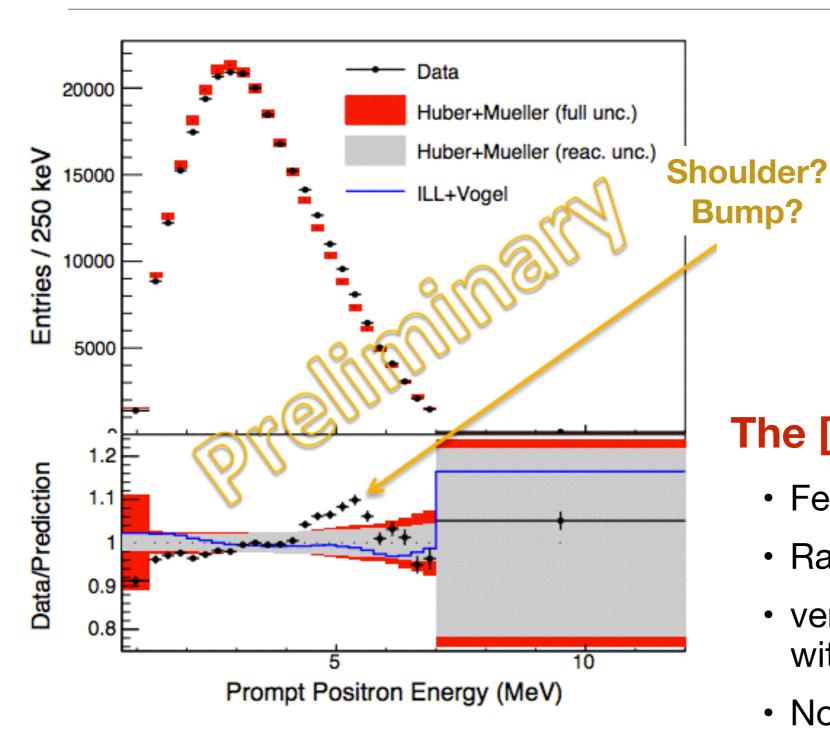


- Two major sources of non-linearity
 - scintillator response: modeled with Birks formula and Cherenkov contribution
 - electronics: modeled with MC and single channel FADC measurement
- Nominal Model: Combined fit with mono-energetic gamma peaks and ¹²B beta-decay spectrum
- Cross-validation Model: ²¹²Bi, ²¹⁴Bi, ²⁰⁸Tl beta-decay spectrum, Michel electron spectrum and standalone bench-top Compton scattering measurement.

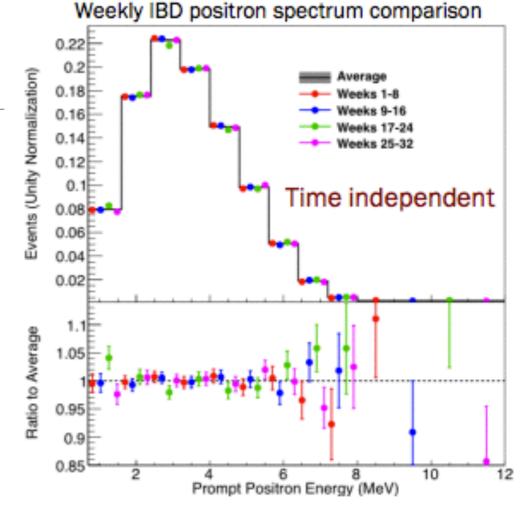


< 1% uncertainty above 2 MeV

Prompt Positron Spectrum



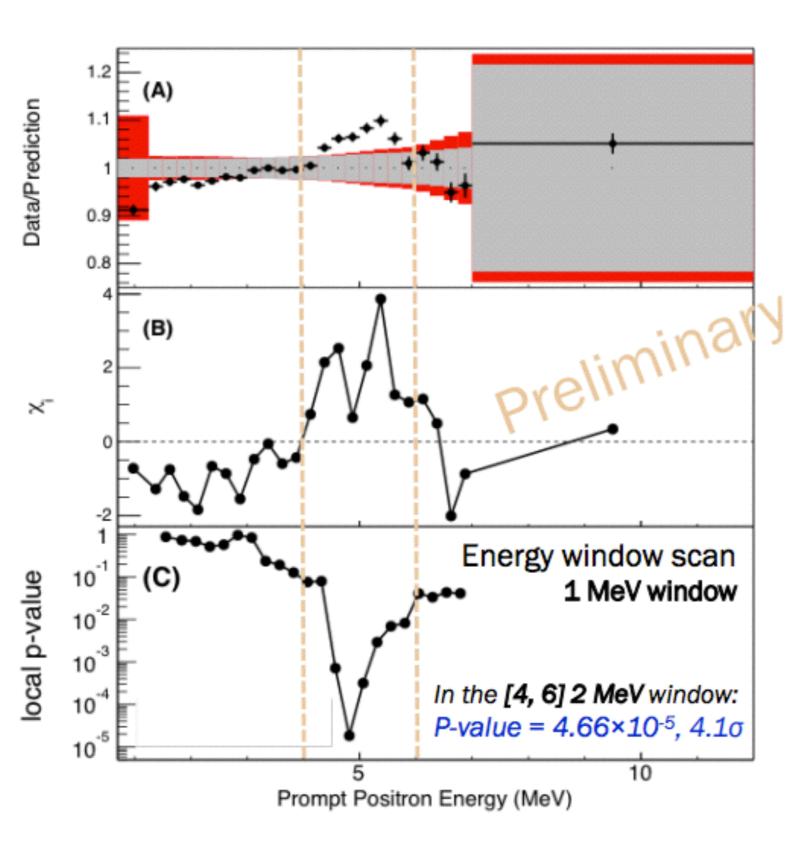
combining the spectrum from the three near-site detectors and compare with theoretical predictions



The [4, 6] MeV "bump"

- Feature is time independent
- Rate correlated with reactor
- vertex/time distribution consistent with IBD
- Not seen in other sample such as 12B beta-decay sample
- Also seen by RENO and DC

Significance of Local Deviation



- (A) Spectral comparison of data and prediction (Huber +Mueller)
 (P-value=0.015, 2.4σ)
- (Β) χ² contribution of each bin, evaluated by:

$$egin{aligned} \widetilde{\chi_i} &= rac{N_i^{ ext{obs}} - N_i^{ ext{pred}}}{\left|N_i^{ ext{obs}} - N_i^{ ext{pred}}
ight|} \sqrt{\sum_j \chi_{ij}^2}, \ where \ \chi_{ij}^2 &= (N_i^{obs} - N_i^{pred}) (V^{-1})_{ij} (N_j^{obs} - N_j^{pred}) \end{aligned}$$

(C) P-value of $\Delta \chi^2$ /ndf in a certain energy window (e.g. 1 MeV)

Introduce N (# of bins) nuisance parameters with no pull terms to oscillation fitter.

Expect the χ^2 difference after introducing the N nuisance parameters follows a χ^2 distribution with N-1 dof.

S(E) [arb.] What causes the "bump"? Data 20000 Huber+Mueller (full unc.) Huber+Mueller (reac. unc.) Entries / 250 keV **Shoulder** 15000 ILL+Vogel **Bump?** 10000 PRL 114, 012502 (2015) 5000 1.2 Data/Prediction 1.1 0.9

combining the spectrum from the three near detectors and compare with theoretical predictions

Prompt Positron Energy (MeV)

- Discrepancy possibly due to uncertainties in predicting the neutrino spectra from the measured beta spectra
 - D. Dwyer, T. Lanford PRL 114, 012502 (2015)
 - A. Hayes et.al, arXiv: 1506.00583

8 decays in the

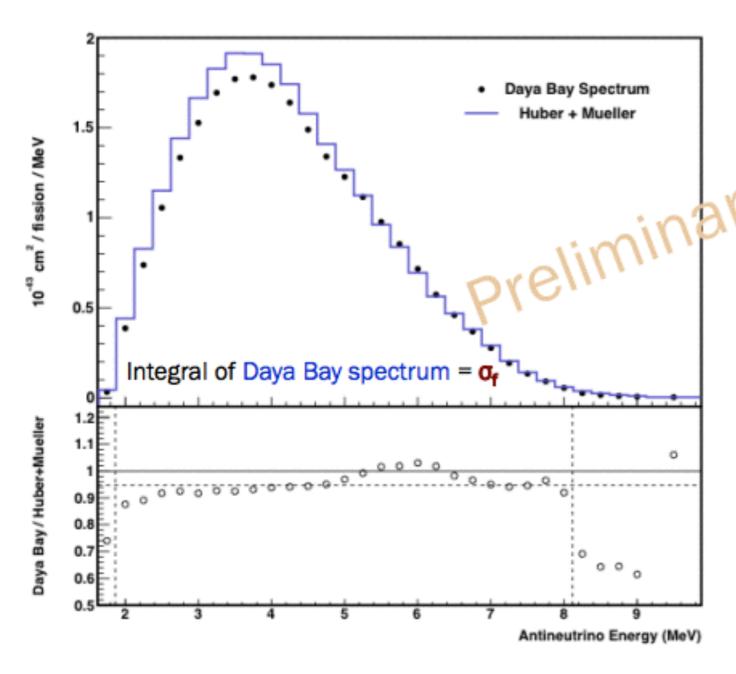
²³⁵U spectrum

dominate the

5-7 MeV (~42%)

Antineutrino Energy [MeV]

Unfolded Reactor Antineutrino Spectrum



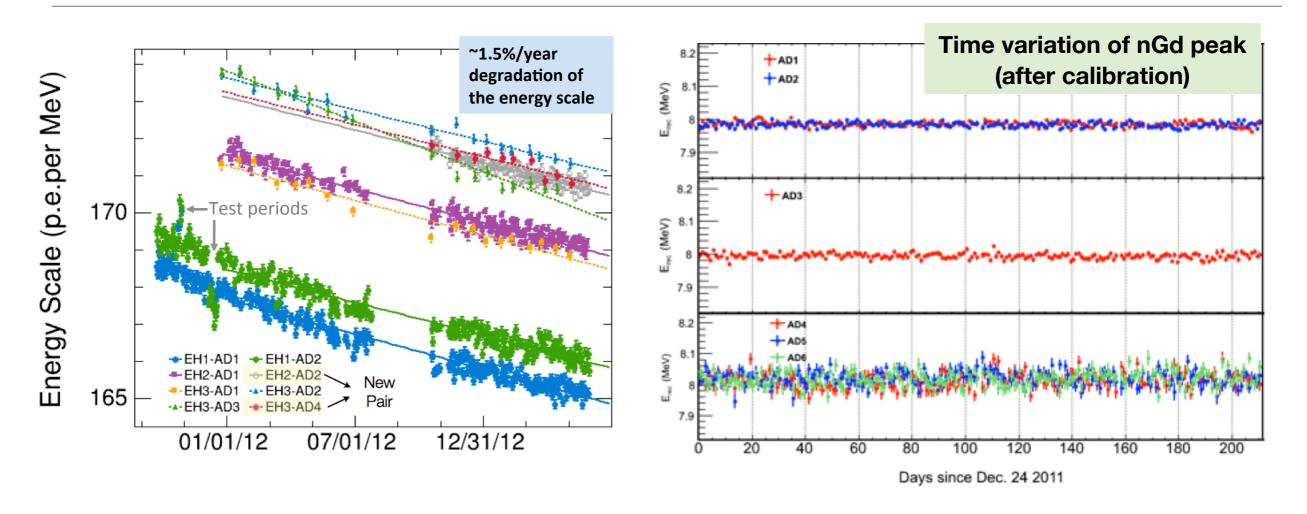
- Unfold the measured positron spectrum into the antineutrino spectrum by removing the detector response and energy resolution effects.
- Can be used directly by future reactor experiments as a reference spectrum
 - more precise than theoretical prediction
 - need small experimentdependent fission fraction corrections

Summary

- Daya Bay updated the oscillation analysis with the full detector configuration
 - Most precise measurement of sin²2θ₁₃ : 6%
 - Most precise measurement of $|\Delta m^2_{32}|$ in the electron neutrino disappearance channel: 4%
- Independent measurement of sin²2θ₁₃ from nH sample
- Set best limit to light sterile neutrinos in 10^{-3} eV² < Δm^2_{41} < 0.1 eV²
- Precision measurement of reactor antineutrino flux and spectrum
 - Flux is consistent with previous short baseline reactor experiments
 - Spectrum has local deviation from prediction in [4, 6] MeV at 4σ level

Backup Slides

Scintillator Stability

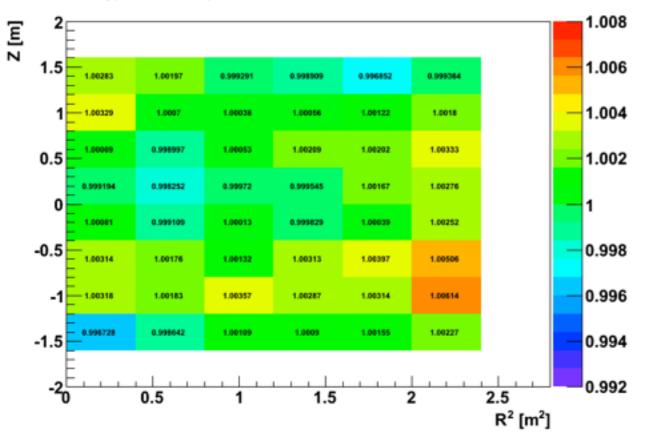


- Small degradation of light yield (~1.5%/year) is seen with nGd, ⁶⁰Co and other event types. Its origin is still unknown (suspect from degradation of attenuation length), but do not anticipate any problems in Daya Bay's lifetime.
- After calibration, achieved energy response that is stable to ~0.1% in all detectors

Energy Uniformity

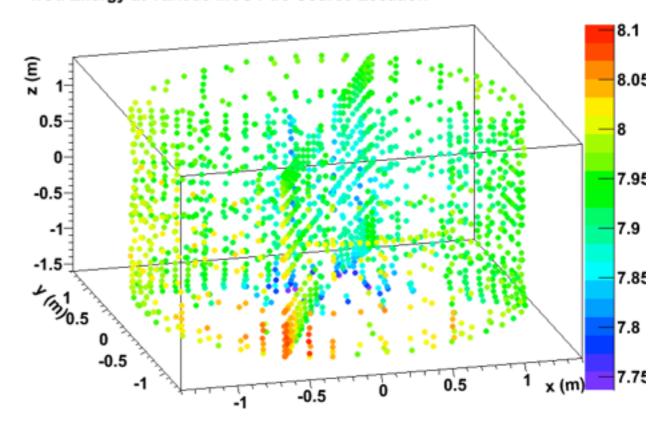
With Spallation Neutron

nGd Energy Uniformity from Spallation Neutron



With MCS Source

nGd Energy at Various MCS PuC Source Location



Inside GdLS volume, < 0.3% variation

Data Table

	EH1		EH2		EH3			
	AD1	AD2	AD3	AD8	AD4	AD5	AD6	AD7
IBD candidates	304459	309354	287098	190046	40956	41203	40677	27419
DAQ live time(days)	565.436	565.436	568.03	378.407	562.451	562.451	562.451	372.685
ε_{μ}	0.8248	0.8218	0.8575	0.8577	0.9811	0.9811	0.9808	0.9811
ε_m	0.9744	0.9748	0.9758	0.9756	0.9756	0.9754	0.9751	0.9758
Accidentals(per day)	8.92 ± 0.09	8.94 ± 0.09	6.76 ± 0.07	6.86 ± 0.07	1.70 ± 0.02	1.59 ± 0.02	1.57 ± 0.02	1.26 ± 0.01
Fast neutron(per AD per day)	0.78 ± 0.12		0.54 ± 0.19		0.05 ± 0.01			
⁹ Li/ ⁸ He(per AD per day)	2.8 ± 1.5		1.7 ± 0.9		0.27 ± 0.14			
Am-C correlated 6-AD(per day)	0.27 ± 0.12	0.25 ± 0.11	0.27 ± 0.12		0.22 ± 0.10	0.21 ± 0.10	0.21 ± 0.09	
Am-C correlated 8-AD(per day)	0.20 ± 0.09	0.21 ± 0.10	0.18 ± 0.08	0.22 ± 0.10	0.06 ± 0.03	0.04 ± 0.02	0.04 ± 0.02	0.07 ± 0.03
13 C(α , n) 16 O(per day)	0.08 ± 0.04	0.07 ± 0.04	0.05 ± 0.03	0.07 ± 0.04	0.05 ± 0.03	0.05 ± 0.03	0.05 ± 0.03	0.05 ± 0.03
IBD rate(per day)	657.18 ± 1.94	670.14 ± 1.95	594.78 ± 1.46	590.81 ± 1.66	73.90 ± 0.41	74.49 ± 0.41	73.58 ± 0.40	75.15 ± 0.49

TABLE I. Summary of signal and backgrounds. Rates are corrected for the muon veto and multiplicity selection efficiencies $\varepsilon_{\mu} \cdot \varepsilon_{m}$. The measured ratio of the IBD rates in AD1 and AD2 (AD3 and AD8 in the 8-AD period) was 0.981 ± 0.004 (1.019 ± 0.004) while the expected ratio was 0.982 (1.012).