

Charge control of an electrically isolated proof mass using a UV LED at 255 nm

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Motivation for DUV LEDs in Space

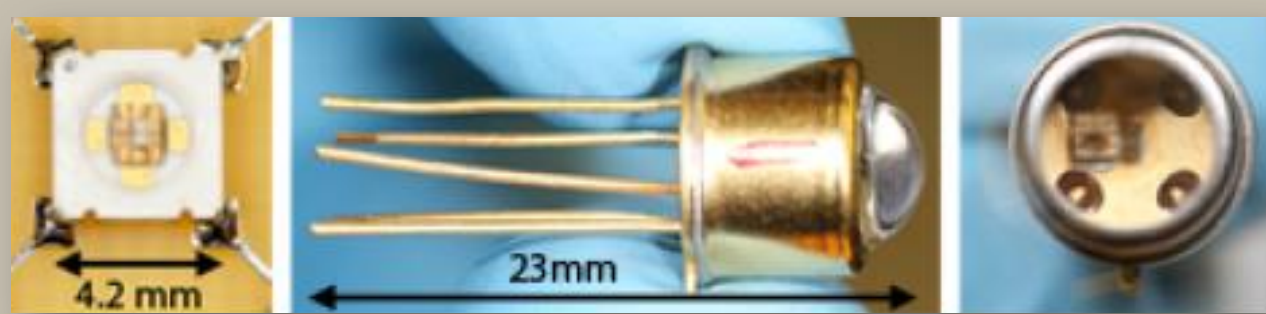
Highly sensitive space-based instruments use drag-free techniques for the control and the science signals

- Laser Interferometer Space Antenna (LISA): $5\text{-}20 \text{ pm/Hz}^{-1/2}$

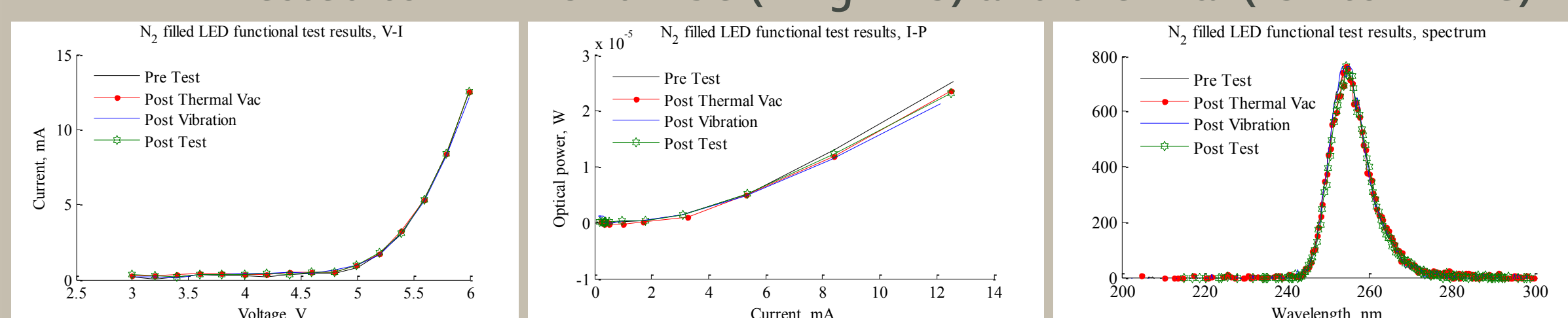
During drag-free flight, the spacecraft follows a free floating proof mass:

- Many disturbances removed
- Charge accumulation on proof mass: $50\text{-}200 \text{ e}^-/\text{sec}$

Use UV photoemission to control charge:



- AlGaIn UV LEDs with 254 nm peak and 12 nm FWHM
- Output power: $\sim 5 \text{ nW}$ to $>100 \text{ }\mu\text{W}$
- Tested to MIL-1540 vibs (14 gRMS) and thermal (-34 to $+71^\circ\text{C}$)



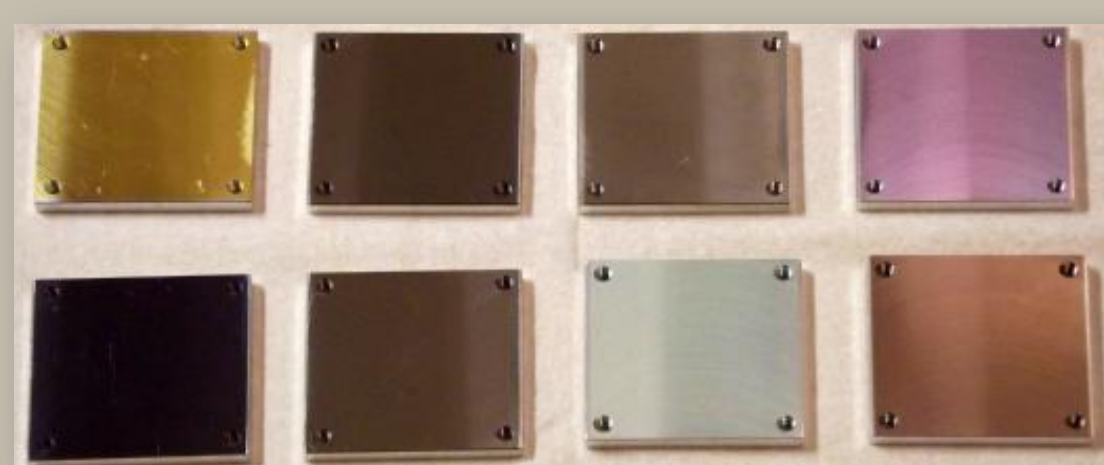
Proof Mass Coatings

Several requirements for coatings:

- Workfunction for UV photoemission ($E_{ph} = 4.86 \text{ eV}$ @ 255 nm)
- Minimize patch effects
- Maintain surface properties during proof mass un/caging
- Minimize reflections from other GRS instruments

Carbide coatings:

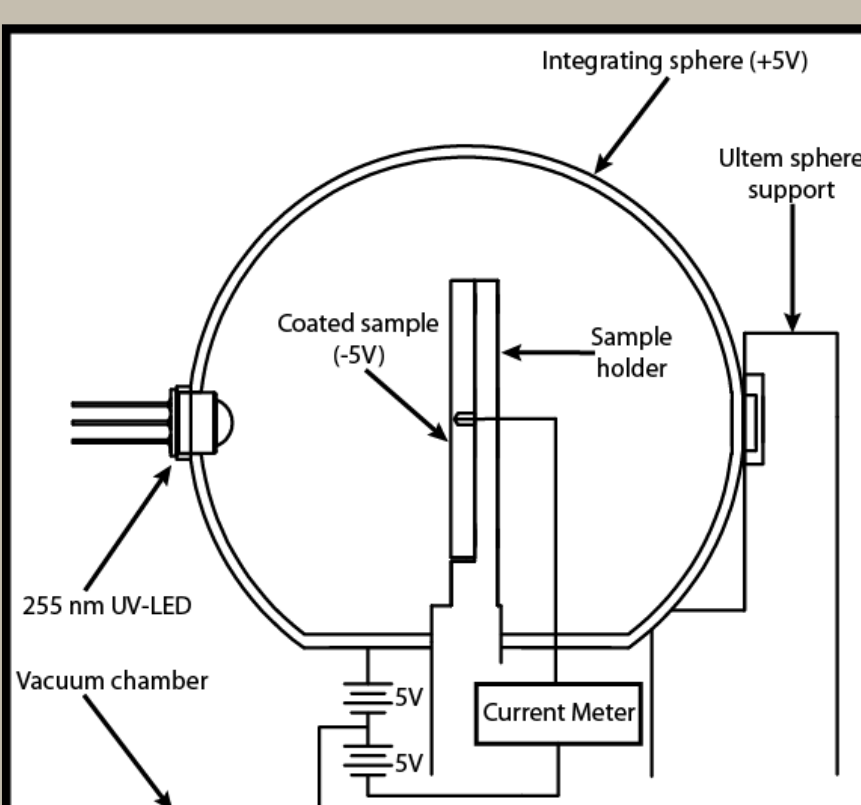
- Workfunctions near 4.86 eV
- High mechanical toughness



Top (L-R): Au, Nb, Ir, SiC.
Bottom (L-R): TiC, Mo₂C, ZrC, TaC

Coating property measurement:

- QE: Keithley 6485 picoammeter, UV LED at $50 \text{ }\mu\text{W}$, 10^{-4} torr
- Reflectivity: UV LED with 45° incident angle, Newport 1931C power meter
- Resistivity: Prometrix Omnimap RS35e with Van der Pauw probe

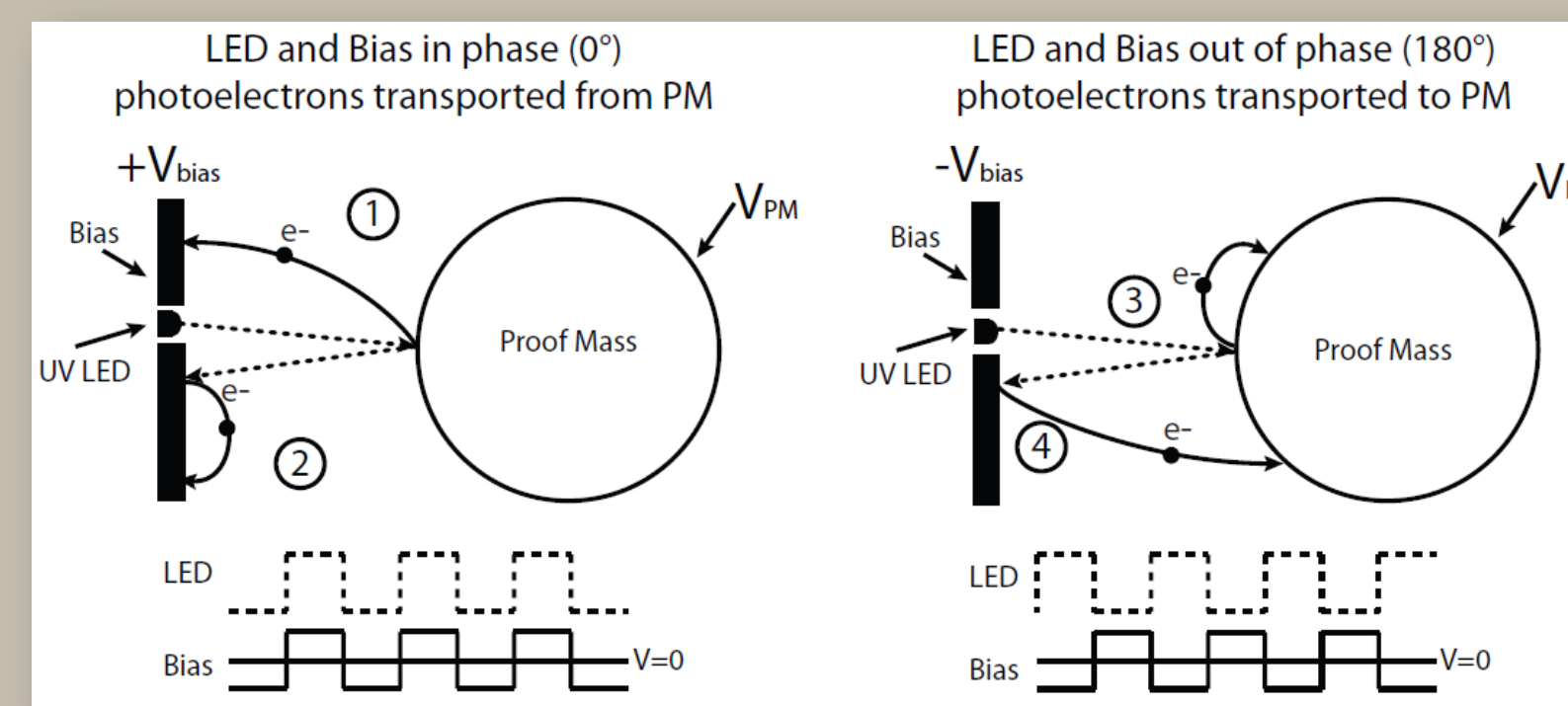


Quantum efficiency measurement test setup

Measured coating properties

Material	Quantum Efficiency	Reflectivity	Resistivity (Ω/sq)
SiC	3.4×10^{-7}	0.12	$2.0 \pm 0.3 \times 10^{-4}$
Mo ₂ C	6.8×10^{-7}	0.15	$1.5 \pm 0.03 \times 10^{-5}$
TaC	6.3×10^{-7}	0.13	$5 \pm 1 \times 10^{-3}$
TiC	4.5×10^{-7}	0.15	$1.16 \pm 0.01 \times 10^{-3}$
ZrC	3.8×10^{-7}	0.11	$7.94 \pm 0.07 \times 10^0$

AC Charge Management Operation



Schematic showing AC Charge Management operation concept

AC charge management uses UV photoemission and two control signals to direct electron flow:

- LED: modulated on/off Bias: modulated $+V_{bias}/-V_{bias}$
- UV light incident on proof mass – some light reflects back to bias plate
- Photoelectrons generated from both surfaces

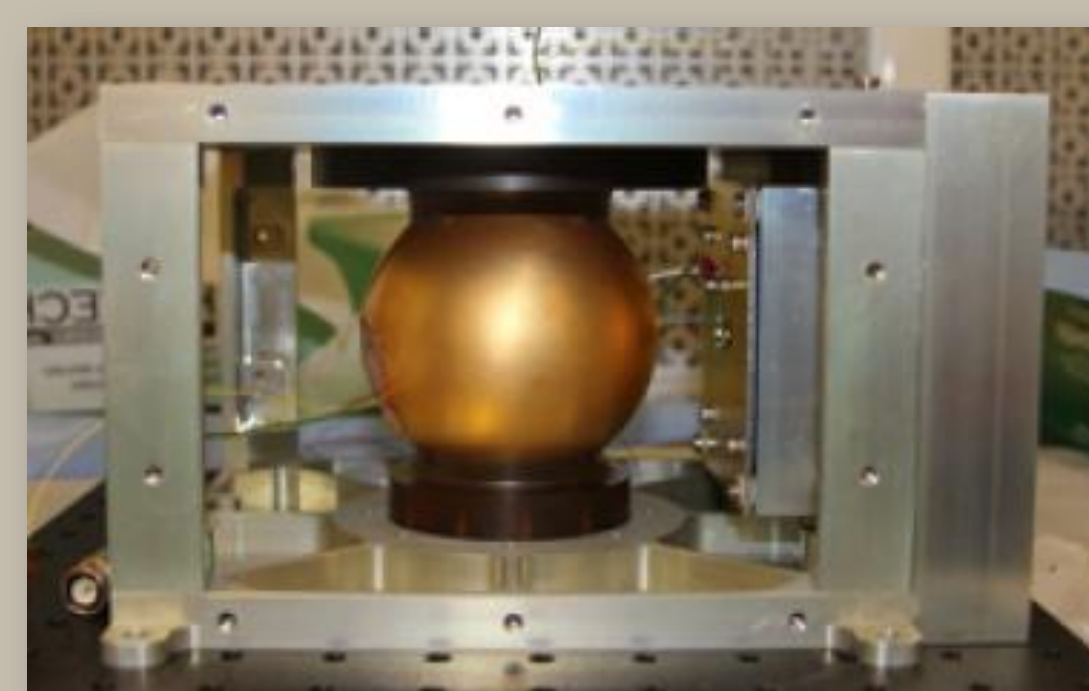
Positive transfer case

1. Bias and LED are in phase – rising edges are synchronized
2. e^- generated when $V_{bias} > 0$
3. e^- pulled away from PM

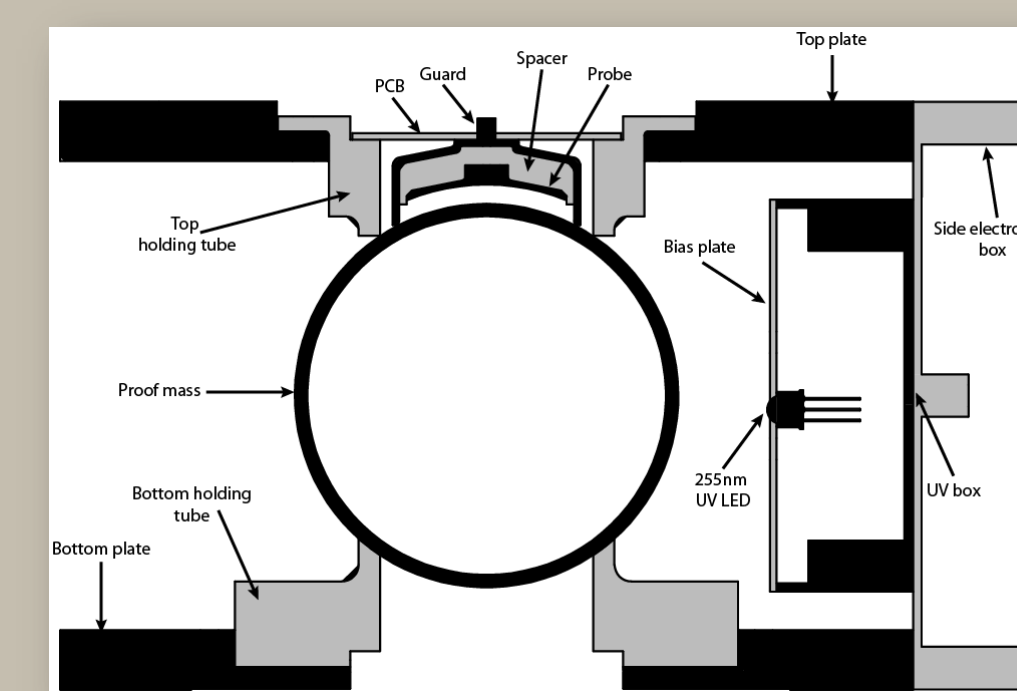
Negative transfer case

1. Bias and LED are out of phase – LED rising/bias falling edges synchronized
2. e^- generated when $V_{bias} < 0$
3. e^- pushed towards PM

AC Charge Management Experiment



Experiment in laboratory configuration



Cutaway showing experiment parts

Demonstration of charge management was performed in a "GRS-like" configuration

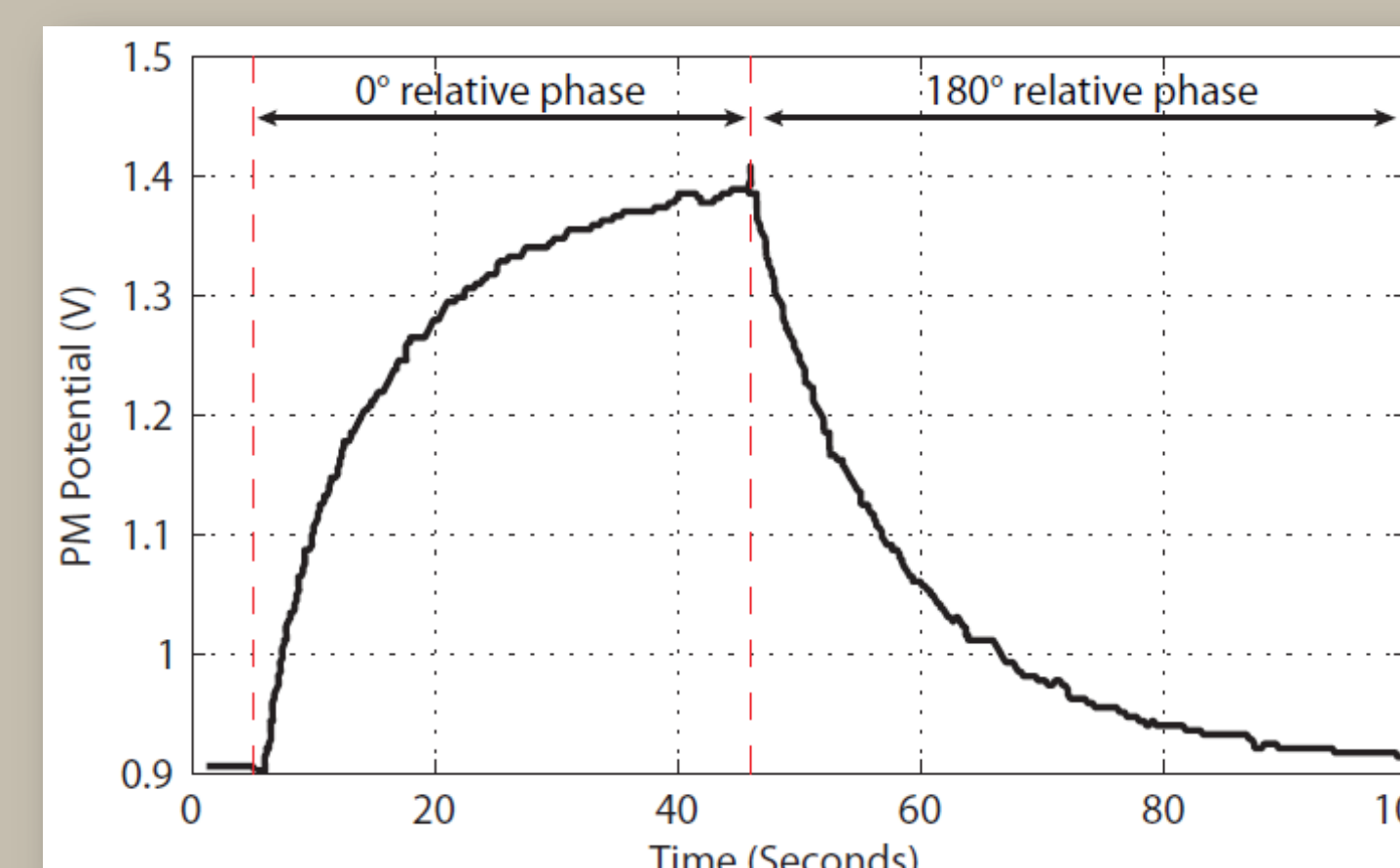
- Single 3.5 in (89 mm) gold plated hollow Al proof mass
- Single UV LED and bias plate
- Sphere supported on Ultem-1000 holding tubes with high R ($>10^{14} \Omega$)
- Potential measured via guarded floating probe held 4 mm from sphere

Experiment parameters

- 10^{-6} torr vacuum
- UV power of $10 \text{ }\mu\text{W}$
- 100 Hz 50% duty cycle
- 3.0 V_{pp} bias
- 17 pF system capacitance

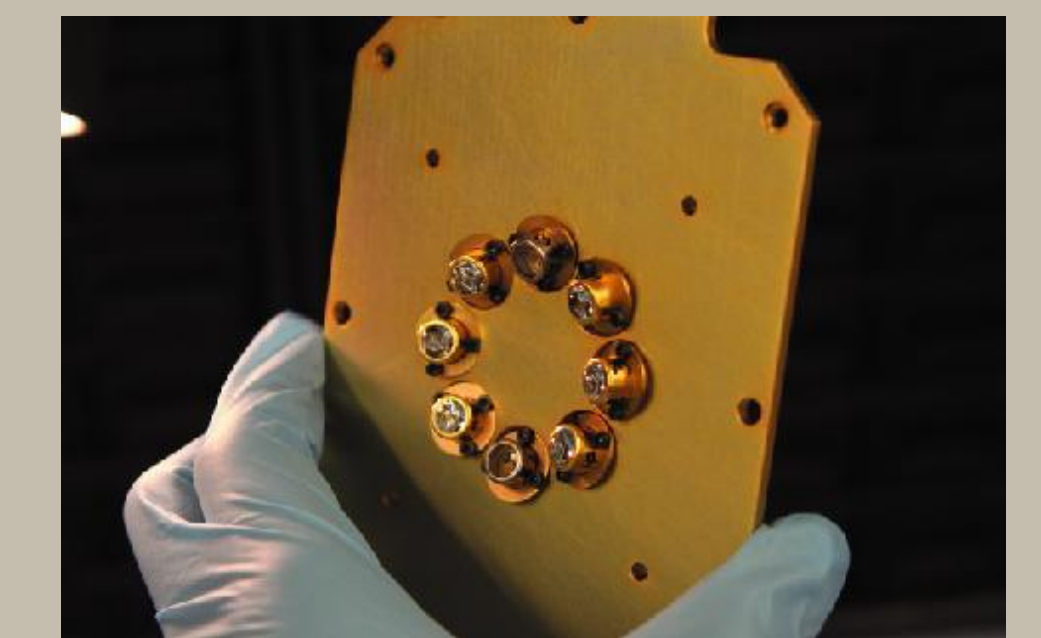
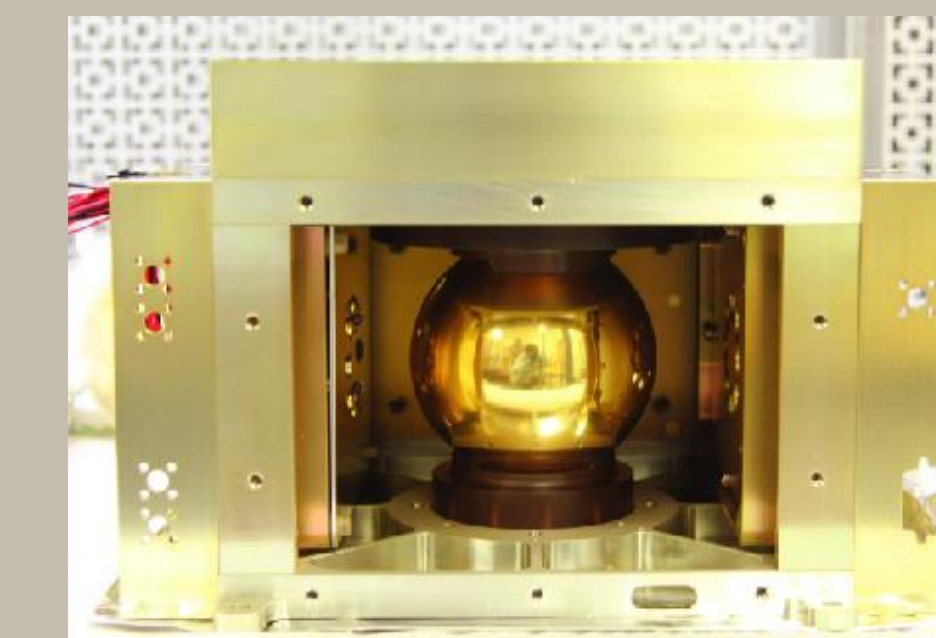
Experiment Results

- Positive: 0.21 pC/sec
- Negative: 0.15 pC/sec



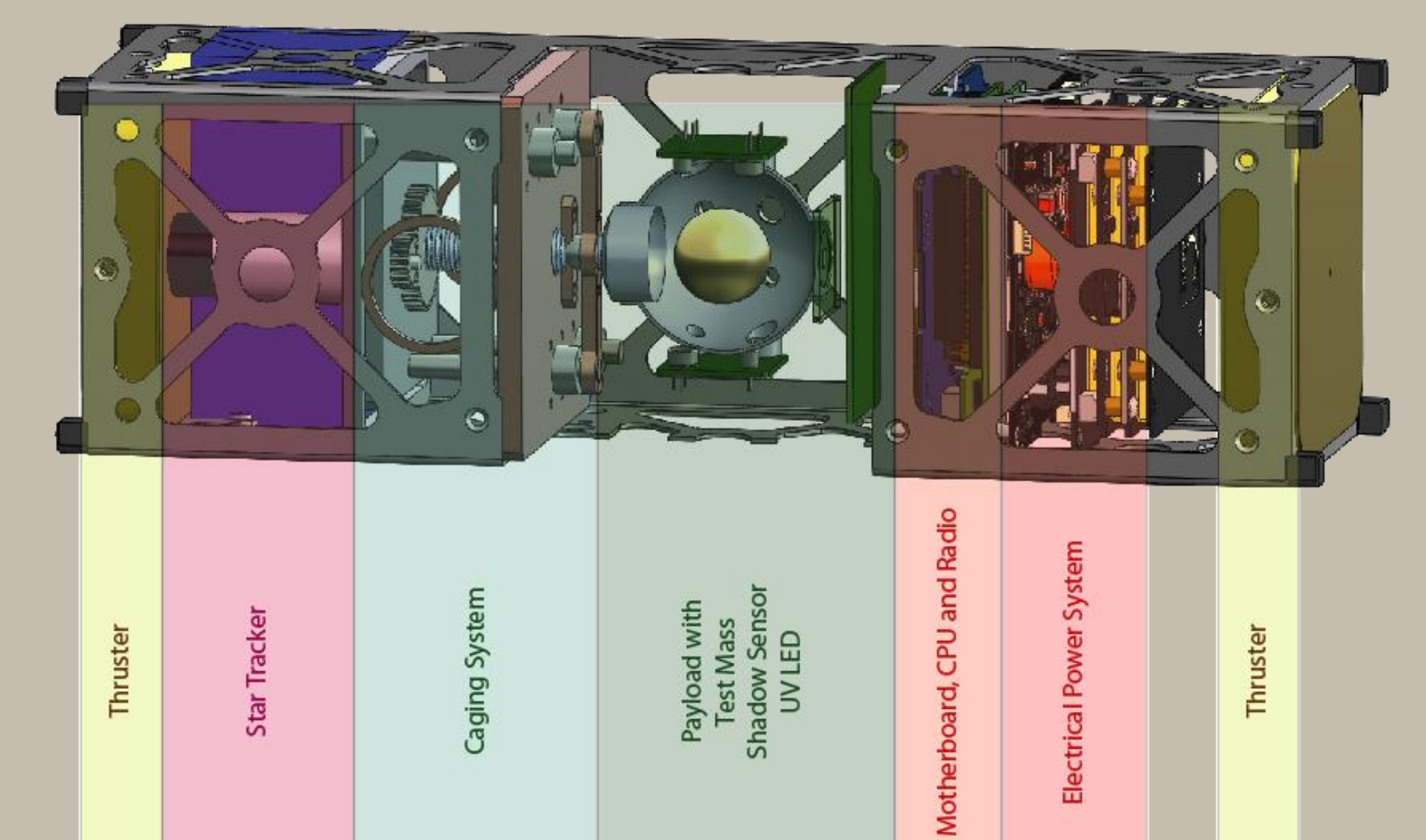
Representative experimental result showing positive and negative charging

Upcoming Missions



UV LED Satellite

- 16 UV LEDs with integrated photodiodes, four bias plates
- Gold coated (deposited) hollow aluminum proof mass
- Raise UV LED and AC charge management system to TRL 8/9
- Joint development with Stanford, NASA Ames and KACST
- Launch in 2013



Drag-free Cubesat

- 3U Cubesat to demonstrate a fully integrated MGRS
- $10^{-12} \text{ ms}^2\text{Hz}^{-1/2}$ from 1 mHz to 1 Hz
- Includes UV LED system for proof mass charge control
- Presented by Andreas Zoellner at the 9th Annual Cubesat Workshop, Cal Poly, San Luis Obispo

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- Mechanical: Dave Muselman (DP Precision)



Please see submitted paper for reference list